



By maths.

To include axisymmetry, define poloidal magnetic flux as:

$$\psi(R,Z) = \int_0^R R' B_z(R',Z) \, dR'$$

And the toroidal current is:

$$-\mu_0 j_\phi = \frac{\partial}{\partial R} \frac{1}{R} \frac{\partial \psi}{\partial R} + \frac{1}{R} \frac{\partial^2 \psi}{\partial Z^2}$$

Going back to terms of Bz:

$$-\mu_0 j_\phi = \frac{\partial B_z}{\partial R} + \frac{1}{R} \frac{\partial^2}{\partial Z^2} \int_0^R R' B_z(R', Z) \, dR'$$

We only see where the MSE emission is, so we can only integrate from some R = RO:



But.... Integral of a second difference of measurement... will be VERY noisy.





So can we directly calculate jphi?



- Predict 30x30 grid of Bz.
- Try to directly calculate j_phi



Conclusion: No. You still cannot exactly calculate jphi directly.

However, we do have measurements of the dBz/dR part at different Zs, and we know that this is most of jphi variation. Together with integral measurements (field pickups and flux loops), it is now part of a complex tomorgraphy problem that we have done before.





By current tomography...

Put description of AUG coils and some pickups into Minerva so we can now do Current Tomorgraphy and Bayesian Equilibrium for AUG.

For magnetics only, we have the usual tomography situation:





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2.69

All sigmaBr = sigmaBz =10mT



By current tomography II

The IMSE still has some a large uncertainty in jphi offset. The unknown term it is not entirely pinned down by the magnetics.

However, the 2D IMSE inference is much better than the equivalent MSE system, for some reason.

Result with Br is much better: If we could get Br as well, we could infer the current almost exactly, within the measurement grid.

Off axis and near the core, the AUG IMSE system will see Br/Bz > 2 with reasonable

Unfortunately, the beam geometry means information about Br is always swamped by Bphi. With NBI **v** in the midplane; $\mathbf{v} \times \mathbf{r}$ and

v × **phi** are always together, regardless of camera view. There is a slight angle though. Full geomtry:

$$\tan \beta \approx \frac{\left(\hat{\underline{v}} \times \hat{\underline{\phi}}\right) \cdot \hat{\underline{r}}}{\left(\hat{\underline{v}} \times \hat{\underline{\phi}}\right) \cdot \hat{\underline{u}}} + \left[\frac{\left(\hat{\underline{v}} \times \hat{\underline{R}}\right) \cdot \hat{\underline{r}}}{\left(\hat{\underline{v}} \times \hat{\underline{\phi}}\right) \cdot \hat{\underline{u}}} - \frac{\left(\hat{\underline{v}} \times \hat{\underline{\phi}}\right) \cdot \hat{\underline{r}}}{\left(\hat{\underline{v}} \times \hat{\underline{\phi}}\right) \cdot \hat{\underline{u}}} \left(\frac{\hat{\underline{v}} \times \hat{\underline{\mu}}}{\hat{\underline{v}} \times \hat{\underline{\phi}}\right) \cdot \hat{\underline{u}}}\right] \frac{B_R}{B_\phi} + \left[\frac{\left(\hat{\underline{v}} \times \hat{\underline{Z}}\right) \cdot \hat{\underline{r}}}{\left(\hat{\underline{v}} \times \hat{\underline{\phi}}\right) \cdot \hat{\underline{u}}} \left(\frac{\hat{\underline{v}} \times \hat{\underline{Z}}\right) \cdot \hat{\underline{u}}}{\left(\hat{\underline{v}} \times \hat{\underline{\phi}}\right) \cdot \hat{\underline{u}}}\right] \frac{B_R}{B_\phi}$$

LOS Intensity averages of coefficients gives:

$$\tan\beta \approx 0.17 + 0.54 \frac{B_Z}{B_\phi} + 0.05 \frac{B_R}{B_\phi}$$

At 5 - 10%, it will have an effect, but we do not expect to see the full current recovery from 2D tomography.

MSE Intensity weighted LOS integral Br/Bz $\int I(l) \frac{B_R}{B_Z} dl / \int I(l) dl$ 0.903 Mag. Axis 🕂 0.903







Para/Diamagnetics

Some notes about Renee's results from the equilibrium point of view:

Just to see, we can load CLISTE's jphi into Minerva and integrate the toroidal flux over the whole vessel (calc. grid). There is a diagmagnetic signal outside the vessel which appears to be uncalibrated. With an offset and scale it mostly agrees with what CLISTE says:



Also, I can now run the code from my PhD work on JET which tries to extract the pedestal pressure from magnetics, wuth the AUG magnetic model. (P. McCarthy has already shown this works at AUG, as I did at JET). With sufficient relaxation of the ff' and p' smoothing priors, it actually finds an equilibrium which is paramagnetic in the very core and diamagnetic at the edge (albeit with a slightly silly pressure profile):



I'm not saying that this is happening, just that with a strong pedestal pressure gradient, it could be.





Poloidal Field at camera

The Minerva AUG magnetostatics model also lets us predict the field from the PF coils:

