



IMSE Design

These slides lay out the design decisions for the ASDEX MSE coherence imaging system. Where possible, this is with first order 'back-of-envelope' calculations so I can get a feeling for the numbers and to cross check the models.

The emphasis is largely on flexibility - so we can cope with errors in input data (e.g. the ray tracing), and to cope, as much as possible, with those pesky 'unknown unknowns'.

I'm not trying to push any boundaries - *yet.* I first want to make it work reliably. The objective of the first phase is to repeatably match the expected polarisation image (or at the very least to a fixed offset) without fudges, hacks or cross calibrations.

The main parameters that need to be decided and the requirements which need to be satisfied are:

Variables:

Focal length and f/# of objective lens.

Focal length and f/# of imaging (camera) lens.

Filter.

Thickness of α BBO delay plate.

Thickness of α BBO displacer and Savart plate.

Requirements:

Image all available FOV of all 4 beams onto CCD.

Set reasonable fringe period (need as much flexibility as possible here!)

Set overall delay to optimise fringe contrast.

Keep as much of the light delivered by forward optics as possible.

Reject as much background spectrum and emit as much useful light as possible.



are currently held - the 'fibre plane'.



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IMSE Design - Imaging



There are some discrepancies with what we think we know:



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We think the fibres are roughly 10x6 grid of 2mm fibres with 1mm core [T. Löbhard].

Ray-tracing of FARO data suggests fibres are ~1.2mm apart vertically.

Original paper [R.Wolf] says 12cm height at beam plane. Here it looks like 18cm.

The final lens L4 has an inconsistent focal length and radius. Radius agrees with imaging, but focal length gets light through PEMs parallel.

However, the 10 channels cover the expected region of the plasma and this matches the ray traced 55cm - 22mm. This is the only important thing, which appears to be consistent.





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IMSE Design - Spectrum

Spectrum across centre of image for high and low fields.





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IMSE Design - Filter.

To look at where is best to filter, plot $|\pi - \sigma| / (\pi + \sigma)$ averaged over scenarios (4 x extreme + 1 middle of the road) This is something like 'generally expected linear polarisation fraction as f(x, λ)':



ldeal filter would almost be sharp high-pass at λ~654.3nm

Interference filter pass-band depends on angle of light, so changes over FOV:

$$\lambda \approx \lambda_0 \sqrt{1 - \frac{1}{2}\sin 2\theta}$$

Above $\theta_{max} \sim 4^{\circ}$ the filter function moves too much to easily capture edges with also capturing poor regions in centre. We really need to keep max angle through filter below 6°.





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IMSE Design - Lenses.







IMSE Design - Lenses.

I've looked around the lab, and around the web for generally available lenses. Zoom (adjustable focal length) lenses tend to not be fast enough for imaging side. We can use one for the objective side though, if it's fast enough and sees the full 24mm virtual image area.

Objective c	andid	ates:	
f	f/#	Req f/#	
75	1.4	2.5	
85	2.1	2.9	
100	1.2	3.3	
17.5 - 105	1.8	3.5@105	
135	2.0	4.5	
180	4.5	6.0	
300	9.0	10.0	

Imaging candidates:		
f	f/#	Req f/#
25	0.85	0.83 🗲
25	0.95	0.83
28	1.4	0.93
35	1.2	1.2
50	1.4	1.6
75	1.4	2.5

L .



We have a box for this.

Some con	nbination	S:				٨c	Ao/Ac		
fo	fo/#	fi	fi/#	Μ	θmax	AC (throughput)	(vignettir	ng)	/57-///////////////////////////////////
75	1.4	25	0.85	33%	9.2°	30%	70%		
75	1.4	25	0.95	33%	9.2°	24% ≬ ⊸	67%		
85	2.1	25	0.95	30%	8.1°	19% l ^a	68%	•	M=100% / 26%
100	1.2	35	1.2	35%	6.9°	17% 🗄	85%	Reasonable	\times
105(Z)	1.8	35	1.2	33%	6.5°	15% 월	86%	options	3
105(Z)	1.8	50	1.4	48%	6.5°	16% ຊິ	87%		
135	2.0	35	1.2	26%	5.1°	9.1%	89%		-10 - 47% ///////////////////////////////////
135	2.0	50	1.4	37%	5 .1°	9.6% g	89% -	Post shaisa of lansas	
180	4.5	50	1.4	28%	3.8°	5.4% <u></u>	92%	(and we have them already)	44200 1/1%
180	4.5	75	1.4	42%	3.8°	5.4%	92%	(and we have cheff an eady)	-15
300	9.0	100	1.2	33%	2.3°	2.0% 🗸	95%		-15 -10 -5 5 10 15

Conclusions:

- Vignetting should not be a problem. Ac/Ae is higher for large θ max, but Ae itself is actually bigger.
- Can change fringe frequency by $\sim 4x$ without changing plates, but at cost of either bad filter shift or low throughput.
- The 180mm/4.5 lens would be really handy, a 35mm/1.2 is almost necessary.
- θ max = 5.1° looks the best middle ground to aim at.

- Throughput for θ max permitted by filter is only 5 - 10%. It is limited by 30mm aperture only for θ max < 5.1°. Increasing crystal size to 35mm aperture would give:

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fo	fi	Ac(30mm)	Ac(35mm)
135	50	9.6% —	→ 13.0%
180	50	5.4% —	→ 7.0%
300	100	2.0% —	→ 2.7%

So bigger plates are not really worth the price.

NB: θ max is the angle through the plates of the light from the edge of visible beam. It isn't exactly equal to the angle of light to the edge of the CCD. i.e. it is independent of M.

Effective CCD size on image at fibre plane



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IMSE Design - Fringes

From F.E.Veiras, phase shift in arbitrary crystal: **NB:** α here is the incidence angle - like θ_{max} was before - sorry.



$$\Delta \phi = \frac{2\pi L}{\lambda_0} \left[\left(n_o^2 - n^2 \sin^2 \alpha \right)^{\frac{1}{2}} + \frac{n}{S} \left(n_o^2 - n_e^2 \right) \sin \theta \cos \theta \cos \delta \sin \alpha - \frac{n_o}{S} \left[n_e^2 S - \left[n_e^2 - \left(n_e^2 - n_o^2 \right) \cos^2 \theta \sin^2 \delta \right] n^2 \sin^2 \alpha \right]^{\frac{1}{2}} \right]$$

Generally, n=1 and $sin^2\alpha$ is small

α

Waveplate (θ =0):

$$\Delta \phi = \frac{2\pi L}{\lambda_0} \left[\left(n_o - n_e \right) - \sin^2 \alpha \, \frac{1}{2n_o} \left(1 - \frac{n_e}{n_o} \left[1 - \sin^2 \delta \left(1 - \frac{n_o^2}{n_e^2} \right) \right] \right) \right]$$

The sin² α term gives the fringes due to the delay plate (which bend the displacer fringes). To quantify, we can calculate α_p , the angle at which it gives 1 full phase rotation:

$$\alpha_p\approx \sqrt{\frac{2n_o\lambda_0}{L\left(1-\frac{n_e}{n_o}\right)}} \quad \begin{array}{l} \text{at } (\mathbf{\delta}=\mathbf{0^\circ},\mathbf{90^\circ}, \mathbf{180^\circ \ or \ 270^\circ} \end{array}$$

45%

This all matches what we see in the lab:



I'll come back to this at the end.

Savart plate (2 displacers at 90°):

$$\Delta \phi_s = \Delta \phi \left(\frac{L}{2}, \delta\right) + \Delta \phi \left(\frac{L}{2}, \delta - \frac{\pi}{2}\right)$$
$$\Delta \phi_s = \frac{2\pi L}{\sqrt{2\lambda_0}} \frac{(n_o^2 - n_e^2)}{(n_o^2 + n_e^2)} \sin\left(\delta + \frac{\pi}{4}\right) \sin\alpha$$

Has no zero-order delay and produces fringes running at 45° to the first plate axis, of the same frequency as a displacer plate of thickness L / $\sqrt{2}$. (i.e. The Savart plate is $\sqrt{2}$ thicker)

For αBBO at 653.5nm:

N = 5.5 fringes per mm per degree

Displacer plate (θ =45°):

$$\Delta \phi = \frac{2\pi L}{\lambda_0} \begin{bmatrix} \frac{(n_o - n_e)}{2} + \frac{(n_o^2 - n_e^2)}{(n_o^2 + n_e^2)} \cos \delta \sin \alpha \end{bmatrix}$$

$$\Box$$
Contribution to fixed delay is ~1/2 of same thickness waveplate +/- 10%

The maximum α is θ_{max} from earlier. So the number of fringes for the full ($2x \ \theta max$) image is:

$$N = \frac{2N\Delta\phi}{2\pi} \approx \frac{2L}{\lambda_0} \frac{(n_o^2 - n_e^2)}{(n_o^2 + n_e^2)} \theta_{max}$$

For α BBO at 653.5nm: $n_o = 1.666, n_e = 1.549$

 $N = 2.2 \times 10^5 L \theta_{max}$ (~

~ 4 fringes per mm per degree



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IMSE Design - Fringes 2

How many fringes do we want? This is the bit I'm least sure about....

When analysing the image, we take the FFT and isolate components. High frequency information moves away from component centres and where they overlap, contaminates the other components. Any space on the high frequency side of the fringe frequency is effectively wasted. So, central frequencies should be at 2/3 of the Nyquist frequency - i.e. 3 pixels per fringe.



320x240, θ_{max} =5.1°, L_{dly} = 4mm L_{sav} = 7.6mm, L_{dsp} = 5.4mm.

However...

- a) We don't yet know what resolution we'll want it will depend on the time resolution and light level. We can say, that the minimum resolution will be limited to 3x the number of fringes.
- b) The highest properly resolvable frequency will be lower than the Nyquist frequency due to imperfect focusing of the fringes by the final imaging lens i.e. the blurring attenuates high frequencies.
- c) Such a high frequency 'feels' like a bad idea, why?

Mathematically, $\Delta \phi$ for the plates (from Veiras) has higher order terms only in $sin^2 \alpha$ and in $(n_e - n_o)$. Since it is all linear in L, there should be no detrimental effects.

Linearity in L also tells us that increasing the fringe frequency by increasing L is better than by increasing θ_{max} .

Max resolution for sensicam, pixelfly etc is 1376x1040. Take 1/4 of that to be very safe and make minimum resolution images easier to handle: 320x240.

Using α BBO at 653.5nm, N_{pixels} =320 and the standard case of θ_{max} = 5.1° (135-50):

Displacer: $L_{dsp} = 5.4$ mm Savart: $L_{sav} = 7.6$ mm	Put these int check all of t	to the full forward mode this - works as expected	l to	
At this thickness, our middle	θ_{max}	~ # fringes across beam	Demodulation resolution at plasma	Min image size
and extreme cases give:	9.2° (75-25)	190	6mm	688x520 (½)
	5.1° (135-50)	100	11mm	344x260 (¼) —
	2.3° (300-100)	50	22mm	172x130 (1⁄8)

Considering that the LOS integration gives us a resolution of \sim 20mm at best, this is good.

θmax was calculated for the edge
of the visible beam, which always covers
55cm of beam, regardless of M.
The # fringes here also used θmax, so is
technically #fringes across the beam, not
necessarily exactly across the image.



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IMSE Design - Delay (Simple Calculation)

The final thing to be decided is the thickness of the delay plate.

Take just the full energy component and the very simple splitting model:

$$λ_0 = 653.5$$
nm
 $Δλ ~ 0.4$ nm (see right)
 w_{λ} : FWHM ~ 0.25nm,
 $\sigma_{\lambda} ~ 0.1$ nm (from spectrum, page 3)

Each gives interference pattern with amplitude:

 $A \propto \exp\left(-rac{\sigma_{\lambda}^2}{2\lambda_0^2}\Delta\phi^2
ight)\cos\Delta\phi(\lambda)$

Phase shift is always $\Delta \phi = \frac{2\pi L}{\lambda} (...)$, so $\Delta \phi(\lambda_0 \pm \Delta \lambda) = \Delta \phi(\lambda_0) \pm \frac{\Delta \lambda}{\lambda_0} \Delta \phi(\lambda_0)$

 $(\lambda_0 - \Delta \lambda)$

Sum of interferograms from each component σ - π^+ - π^- gives:

$$A \propto \exp\left(-\frac{\sigma_{\lambda}^{2}}{2\lambda_{0}^{2}}\Delta\phi^{2}\right) \left[1 - \cos\left(\frac{\Delta\lambda}{\lambda_{0}}\Delta\phi\right)\right] \cos\Delta\phi$$

This thing is the 'net contrast' and varies slowly with $\Delta \phi$ i.e. only the zero order $\Delta \phi$ terms matter (no *sin* α)

W

 λ_0

 $(\lambda_0 + \Delta \lambda)$

Maximum is at $L \sim 4.1mm$ for a single delay plate. At the field extremes the max is L = 4.8mm at B = 1.9T and 3.6mm at B = 2.7T.

We already have a $L_{dsp} = 5.4mm$ displacer plate, which gives an effective delay of a 2.7mm delay plate. With the (surface projected) optic axes aligned, the delays add and the delay plate needs to be only 1.4mm. With the plates at 90°, the delay subtracts and the delay plate must be 6.8mm.

From 2 slides ago, the angle at which the 2nd order delay plate terms start to really bend the fringes α_p :

 $\alpha_p \approx \sqrt{\frac{2n_o\lambda_0}{L\left(1-\frac{n_e}{n_o}\right)}} \qquad \text{For } \mathsf{L}_{\mathsf{dly}} = 1.4 \text{mm: } \alpha_{\mathsf{p}} = 8.5^\circ \text{ - nicely outside our 5.1}^\circ \text{ standard case.} \\ \text{For } \mathsf{L}_{\mathsf{dly}} = 6.8 \text{mm: } \alpha_{\mathsf{p}} = 3.9^\circ \text{ - things start to bend.} \end{cases}$

We should use the thinner delay plate and align it with the displacer.

Splitting, very roughly.... B = 2.3T (1.9 to 2.7 is range) Full beam energy = 60 keV (Deuterium = 2 x m_p) Angle between B and $v \sim 62^{\circ}$ in centre. $\lambda_0 = 653.5nm$

$$\begin{split} v &= \sqrt{\left(\frac{2 \times 60 keV}{2m_p}\right)} \approx 2.4 \times 10^6 m s^{-1} \\ E &= \underline{v} \times \underline{B} \approx 4.9 \times 10^6 V m^{-1} \\ \Delta E &= 3 \times \frac{3}{2} e a_0 |E| \approx 1.3 \times 10^{-3} eV \\ \Delta \lambda &= \frac{\lambda_0^2}{hc} \Delta E \approx 0.4 nm \end{split}$$

Ranges from 0.32nm at 1.9T to 0.46nm at 2.7T.







IMSE Design - Delay (Modelling)

That delay calculation might be a bit *too* crude, since the spectrum is considerably more complex, it changes over the FOV with different beams, we need to include dispersion etc.

The virtual camera model (the medium level one) works by adding the DSH equation for each stokes spectral component vector:



We calculate the net spectral contrast C, by building 4 images of the oscillating terms with shifted phases and adding them in quadrature:

 $C^{2} = A(\omega_{0}\tau_{1}, \omega_{0}\tau_{2})^{2} + A(\omega_{0}\tau_{1} + \frac{\pi}{2}, \omega_{0}\tau_{2})^{2} + A(\omega_{0}\tau_{1}, \omega_{0}\tau_{2} + \frac{\pi}{2})^{2} + A(\omega_{0}\tau_{1} + \frac{\pi}{2}, \omega_{0}\tau_{2} + \frac{\pi}{2})^{2}$



We can also plot the optimum L over the 2D image, which shows that it changes more up/down the image:



Optimal here is slightly lower than simple calculation, at 1.2mm instead of 1.4mm. The region of high contrast is broad and doesn't change much between configurations. Changes to the filters, which beams are in use, etc don't really make a big difference either.

Best all round average:
$L_{dly} = 1.2 mm$







IMSE Design - Summary

50mm f/1.4

Variables:

Focal length and f/# of objective lens..... 135mm f/2.0

Focal length and f/# of imaging (camera) lens....

Filter.....

Thickness of αBBO displacer and Savart plates...

Thickness of α BBO delay plate..... $L_{dly} = 1.2mm$

Requirements:

Image all available FOV of all 4 beams onto CCD.....

Set reasonable fringe period (need as much flexibility as possible here!).....

Set overall delay to optimise fringe contrast.....

Keep as much of the light delivered by forward...... optics as possible.

Reject as much background spectrum and emit as...... much useful light as possible. $\theta_{max} = 5.1^{\circ}$ M = 37%

 $\lambda_{max} = 654.5 nm$, tilted at 1° and > 2nm width.

 $L_{dsp} = 5.4 mm, \ L_{sav} = 7.6 mm$

Can in principal range from 50 to 190 fringes.

Easy, this is pretty insensitive.

Only about 10% in standard case, we can try pushing it up to 30% later, but we might have to do something different with the filter.

If we can get the filter flat, this is quite good.

Things to come later: Alternative filter positions, Absolute light at camera, Exposure time



5.1° (135-35)



IMSE Design (Additional) - Alternative configs for filter?

What about a 3+ lens system for dealing with the filter angle so that we can use the higher throughput lens configs?



Try replacing imaging lens with: Imaging - Field - Imaging combination ...



Each pixel now goes through filter at a range of angles - but average still varies too much.

Try using long local focal length first imager and set field lens to get all central ray of each pixel parallel to axis ...



Maybe not, this is getting silly.

Interestingly, for the strongest config (θ max = 9.2° with 75-25), the 25/0.95 imaging lens vignets the light enough for all pixels to have the same Separation of Smallest Plates Smallest Plates average angle at the fibre 35mm 30mm diameter plane, so the filter can just go there instead: Filter

Vignetting is now Ae/Ac = 67%, but since throughput is now Ac=24%, which is 2.5x in 5.1° standard case, the edge light is Ae=2x the 5.1° case.

So it is worth modelling what putting the filter in an image plane actually does.

* Need to work out for the general, or at least the 100/35 case (6.9°) , what field lens at the fibre plane is required to achieve the same result.