



# Motional Stark Effect Imaging on ASDEX Upgrade: 2012 / 2013 Results

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**November 2012:** System assembled and tested with full MSE simulation in HGW lab.

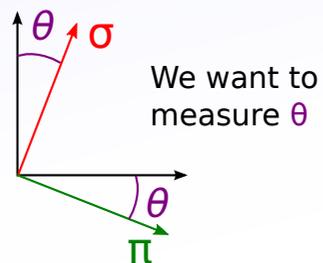
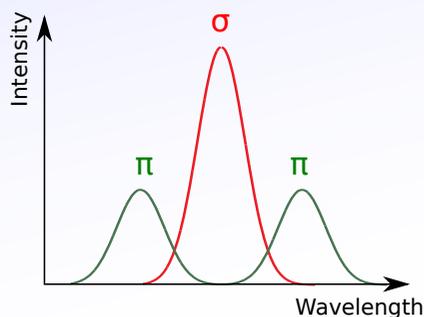
**January 2013:** Installed at AUG and recorded first results - 5 days of plasmas with Deuterium NBI.

**April 2013:** Reinstalled at AUG with various improvements. Hydrogen NBI.

# Imaging Motional Stark Effect at AUG

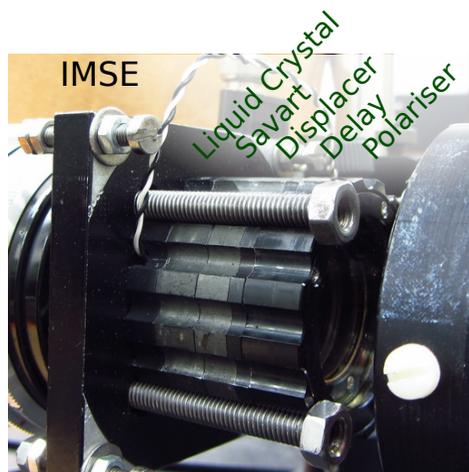
The IMSE system fits to the back of the the existing MSE optics.

H $\alpha$ /D $\alpha$  beam emission is Doppler shifted by beam velocity and split by the Motional Stark Effect into  $\pi$  and  $\sigma$  components which are polarised perp/parallel to projected  $\mathbf{v} \times \mathbf{B}$  direction.

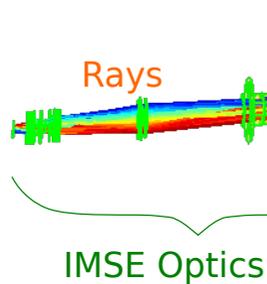
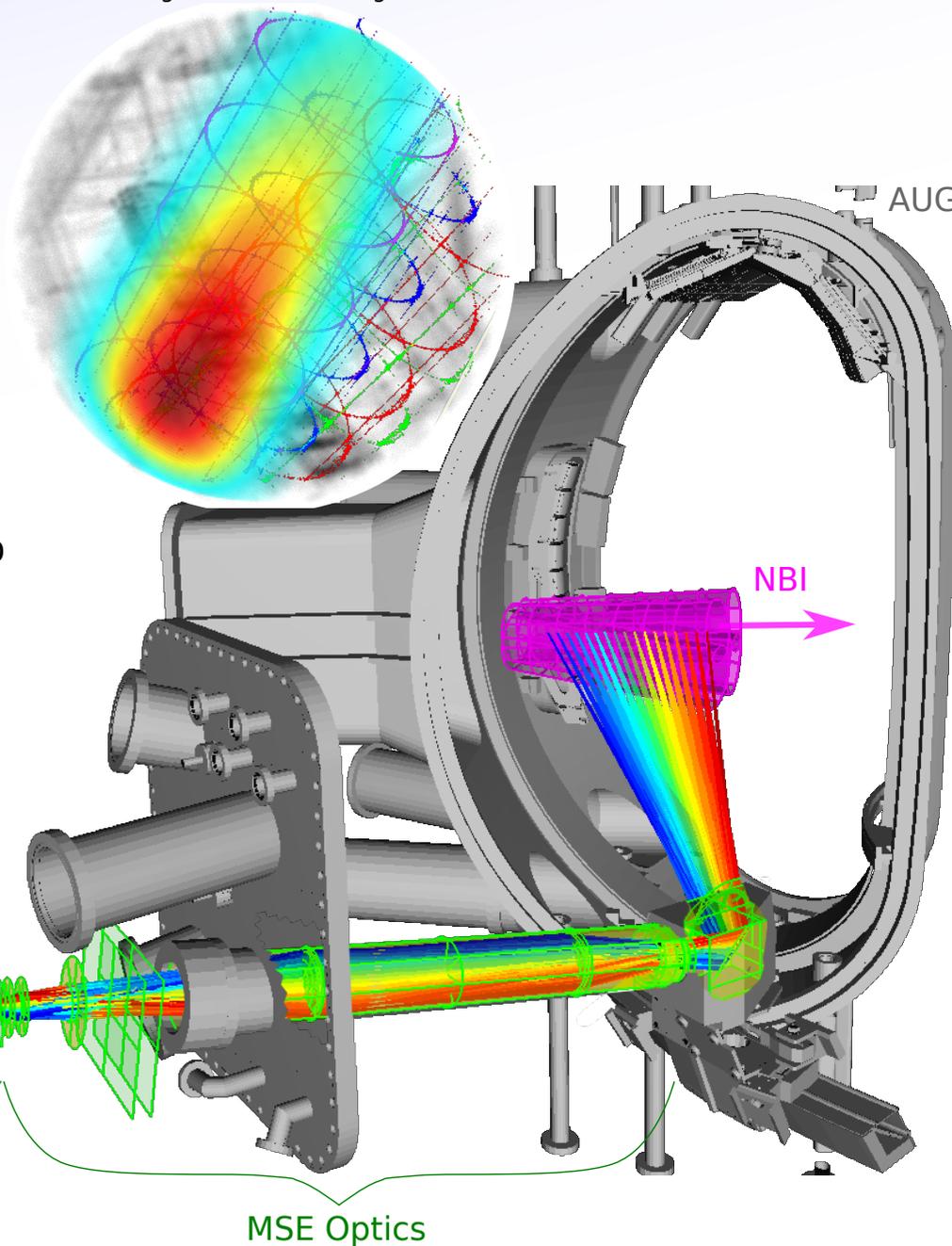


The IMSE records a full image of the beam emission with a CCD camera. A series of birefringent plates superimpose an interference pattern containing the polarisation information.

The difference in the spectrum of the  $\sigma$  and  $\pi$  is exploited to cause their interference patterns to add and the measurement of  $\theta$  is periodic in  $90^\circ$ , so the  $\sigma$  and  $\pi$  directions are averaged.



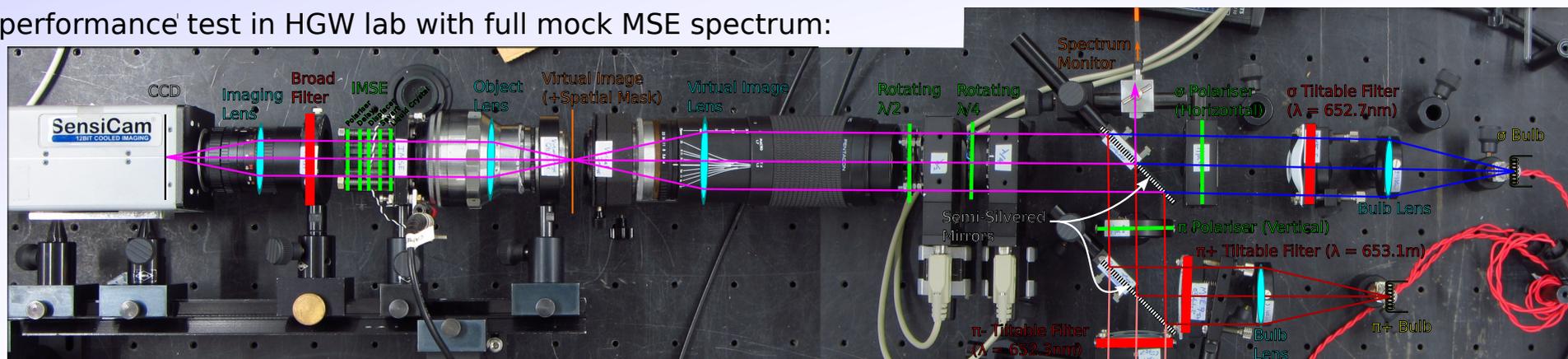
Rendering of beams + background



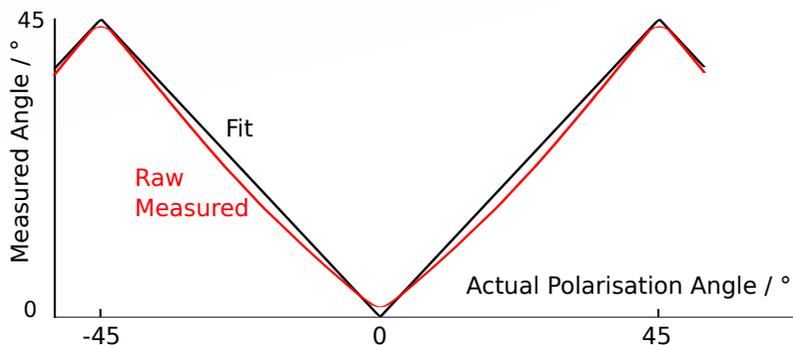
MSE Optics

# November 2012: Performance Tests.

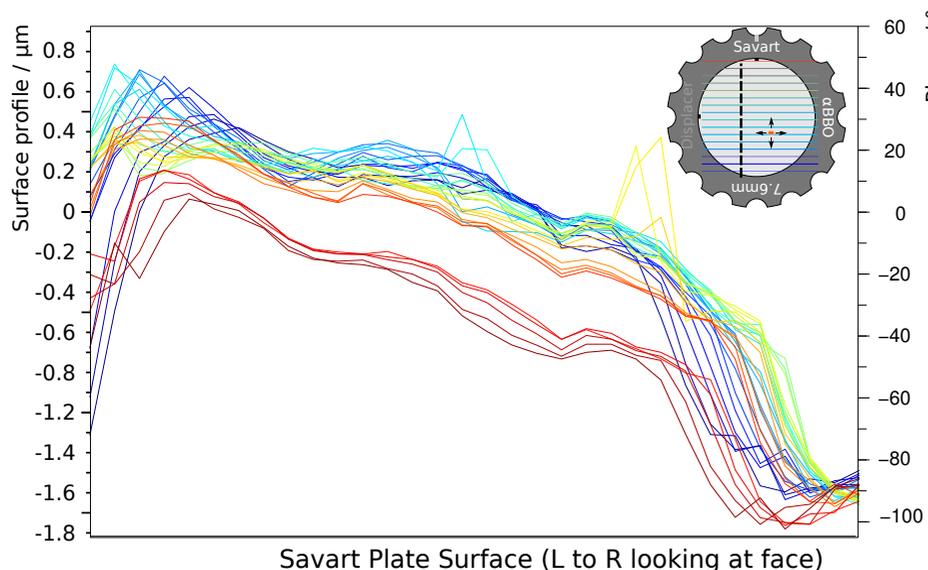
Concept and performance test in HGW lab with full mock MSE spectrum:



Tests revealed a non-linearity in the polarisation measurement:

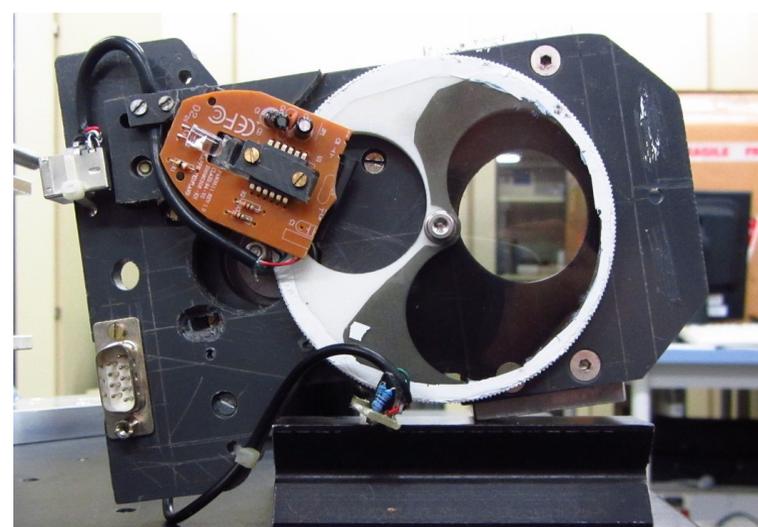


A long investigation found this to be caused by large variations in the Savart plate surface profile:



It depends on input light cone, so requires calibration with the actual beam light. Solution was to quickly build a polariser calibration wheel:

Tests were otherwise successful and proved accuracy down to the required 0.3°.



# January: First installation

Diagnostic assembled and fitted to ASDEX Upgrade for 5 shot days over two weeks.  
Temporarily replaced the existing MSE.

## MSE

Per Channel:

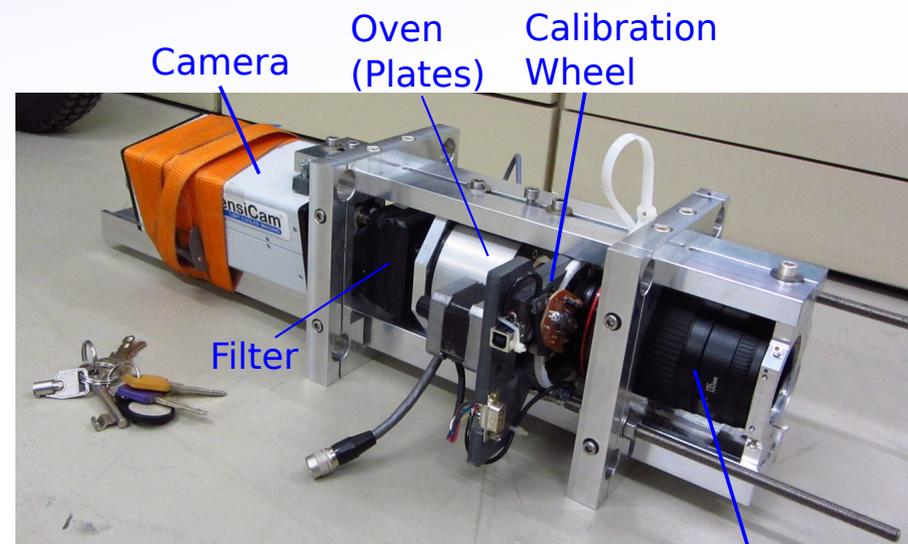
- Optic fibre from torus hall.
- Narrow spectral filter.
- Filter optics
- Oven
- Photo-diode/PMT

PEMs

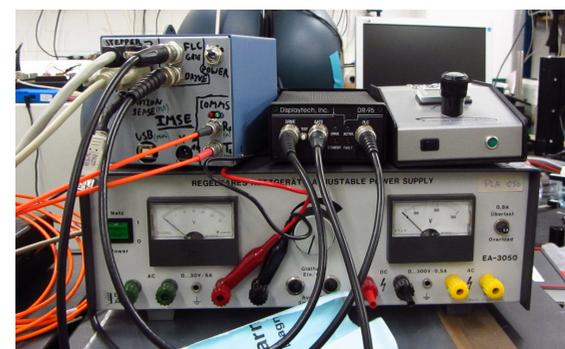
Control rack:

- PEMs controllers
- Lock-in amplifiers (or fast ADCs)
- Control PCs

## IMSE

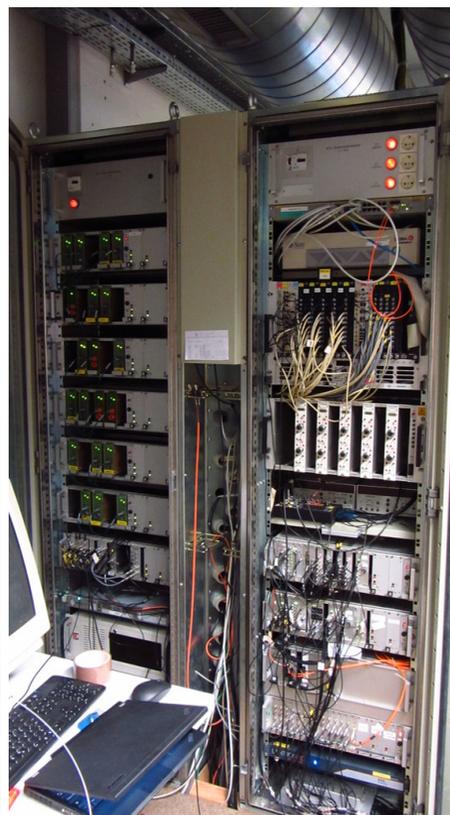
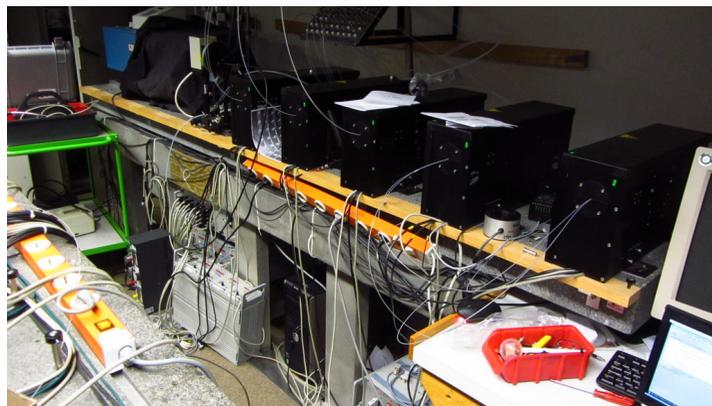


Object Lens



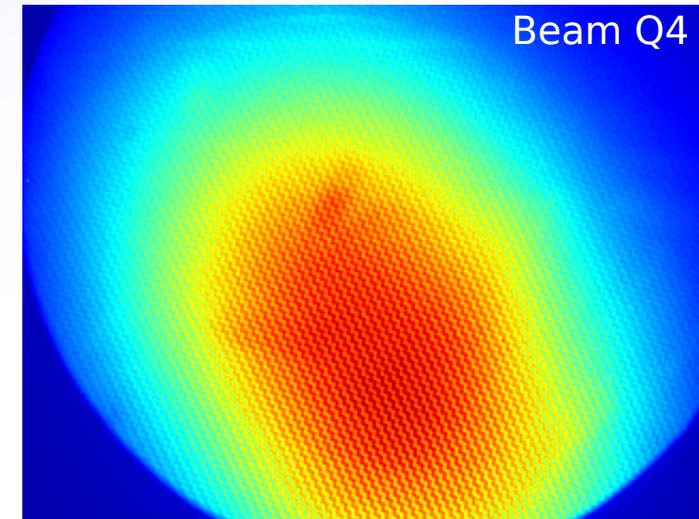
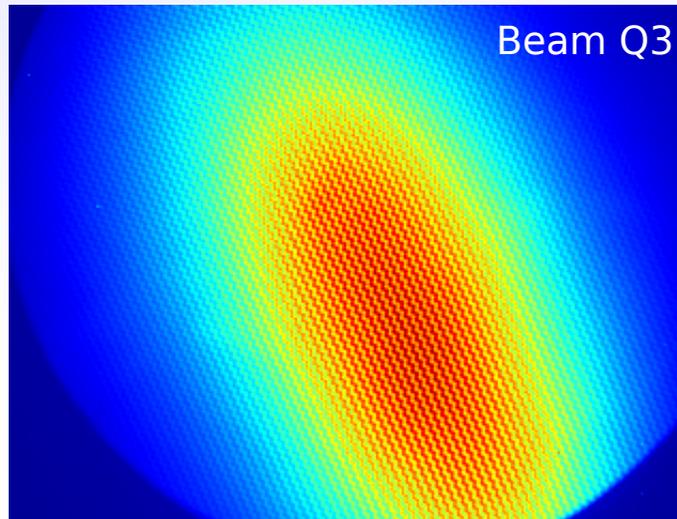
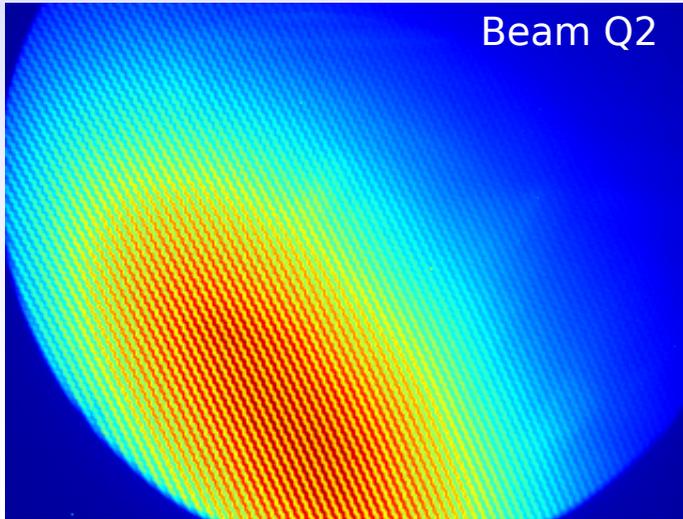
Torus hall controller (40€ Arduino) + power supplies.

Control PC:  
Camera card  
ADC card.

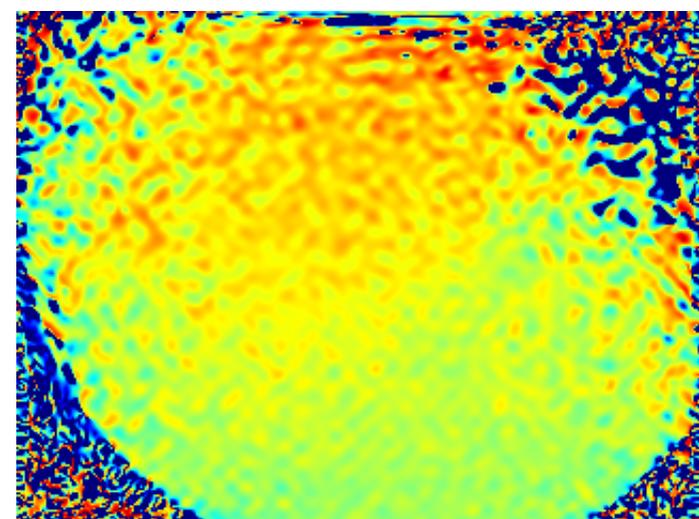
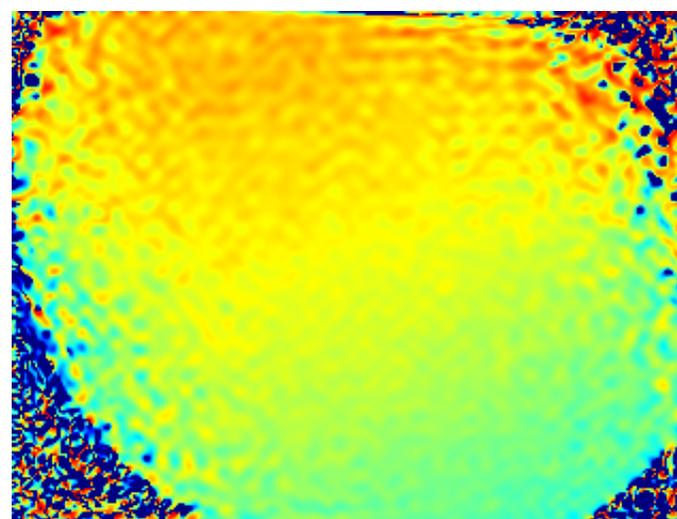
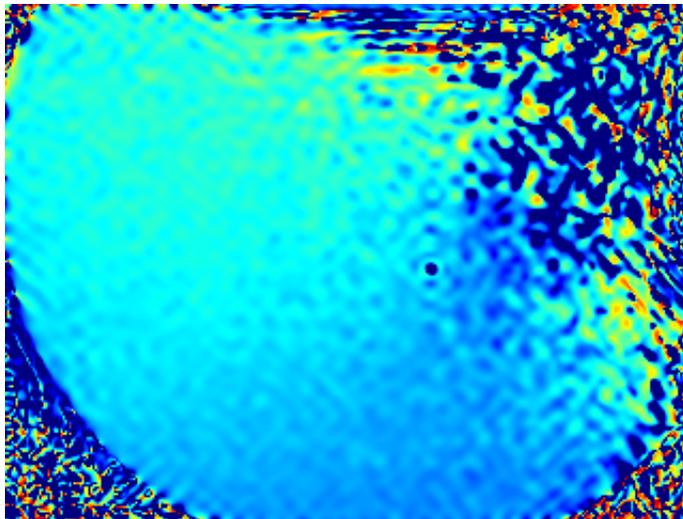


# January: First images

Raw images from first day of operation: Beam emission with interference pattern:

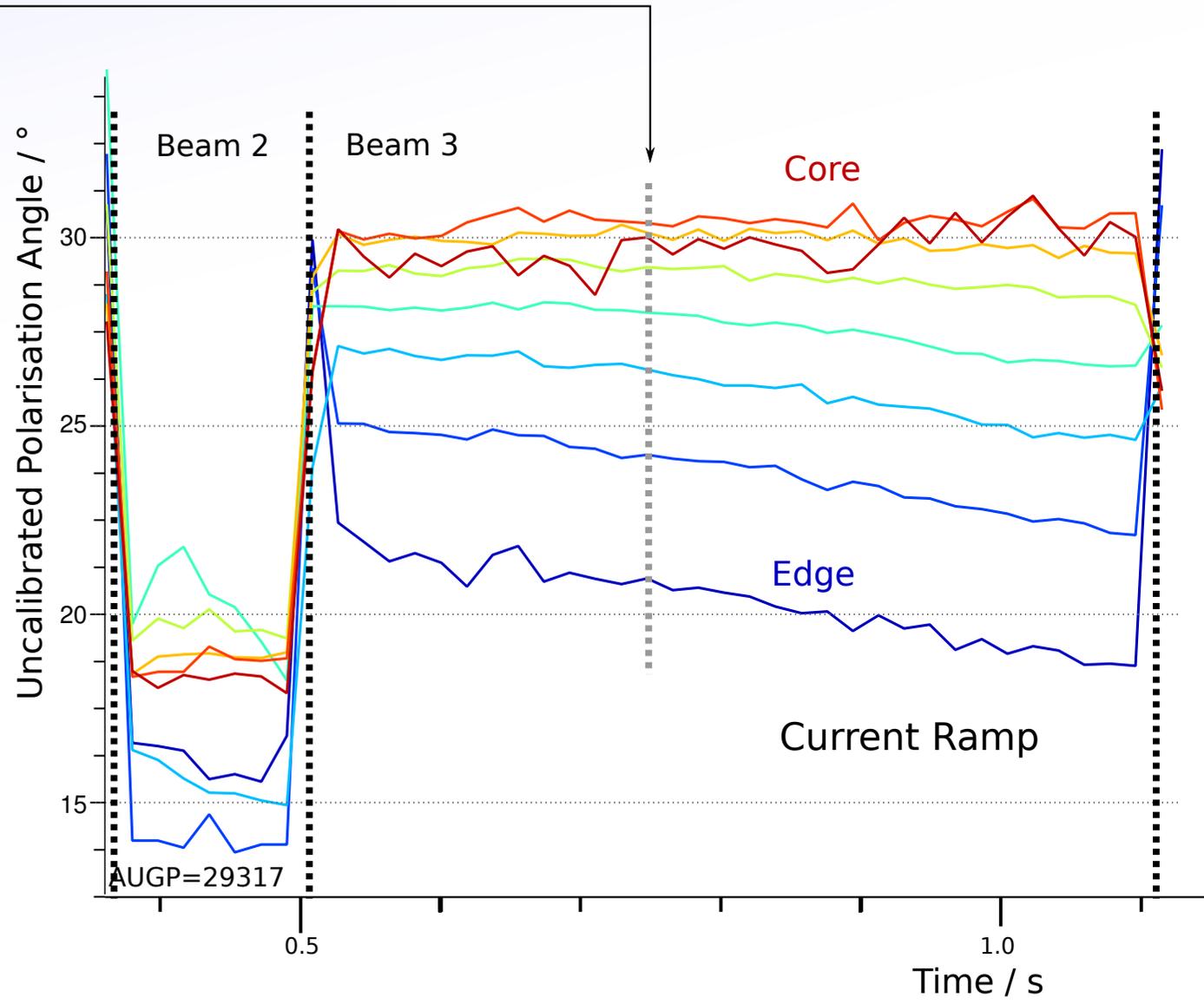
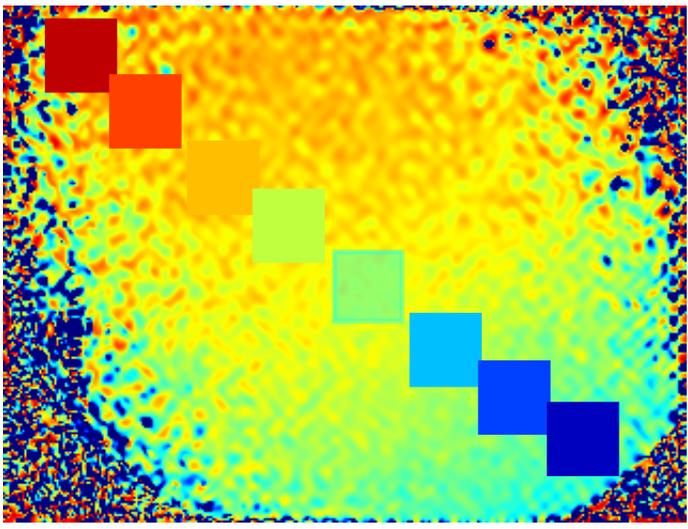
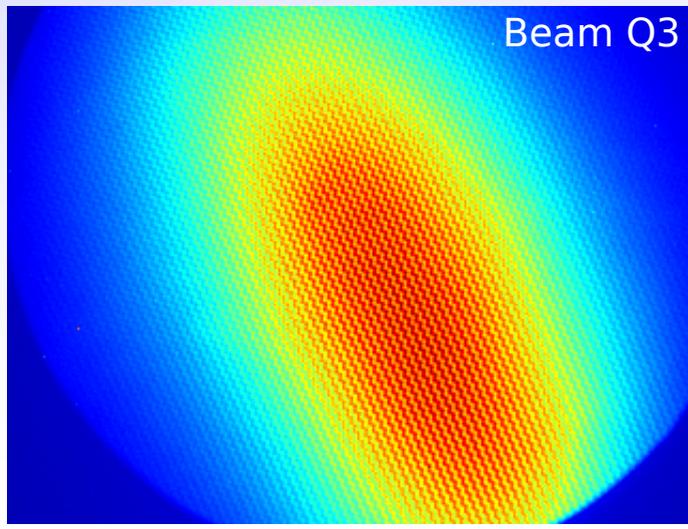


Raw demodulated polarisation angle:



# January: First indication of success:

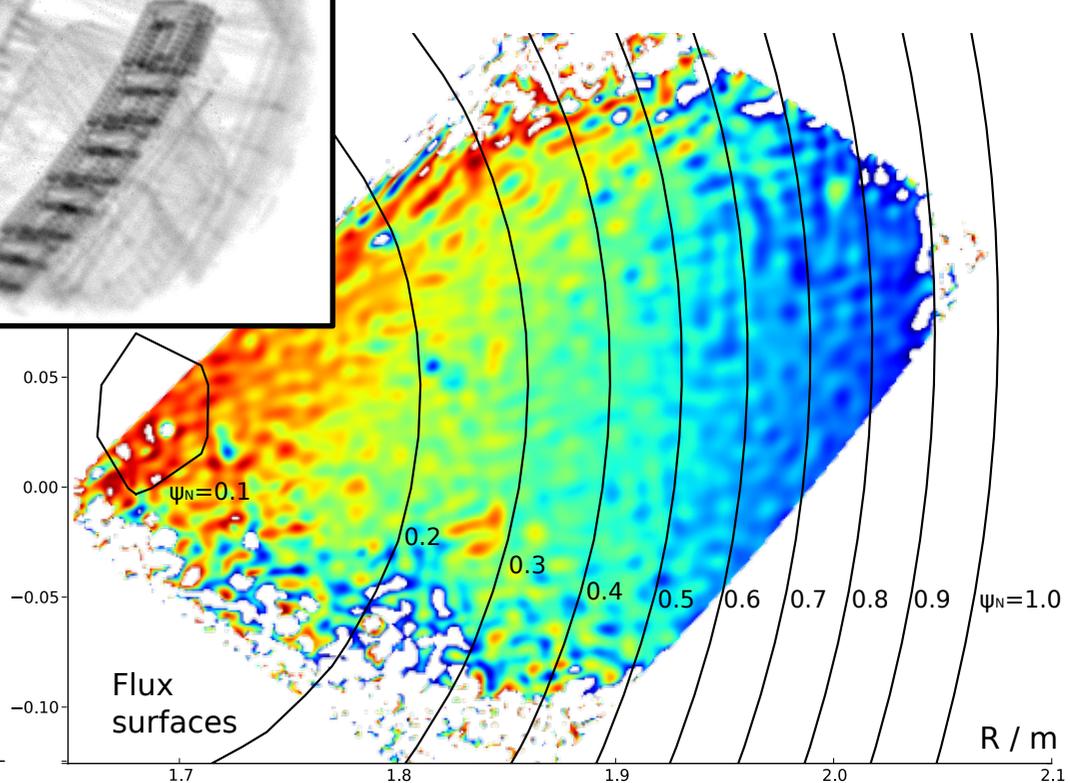
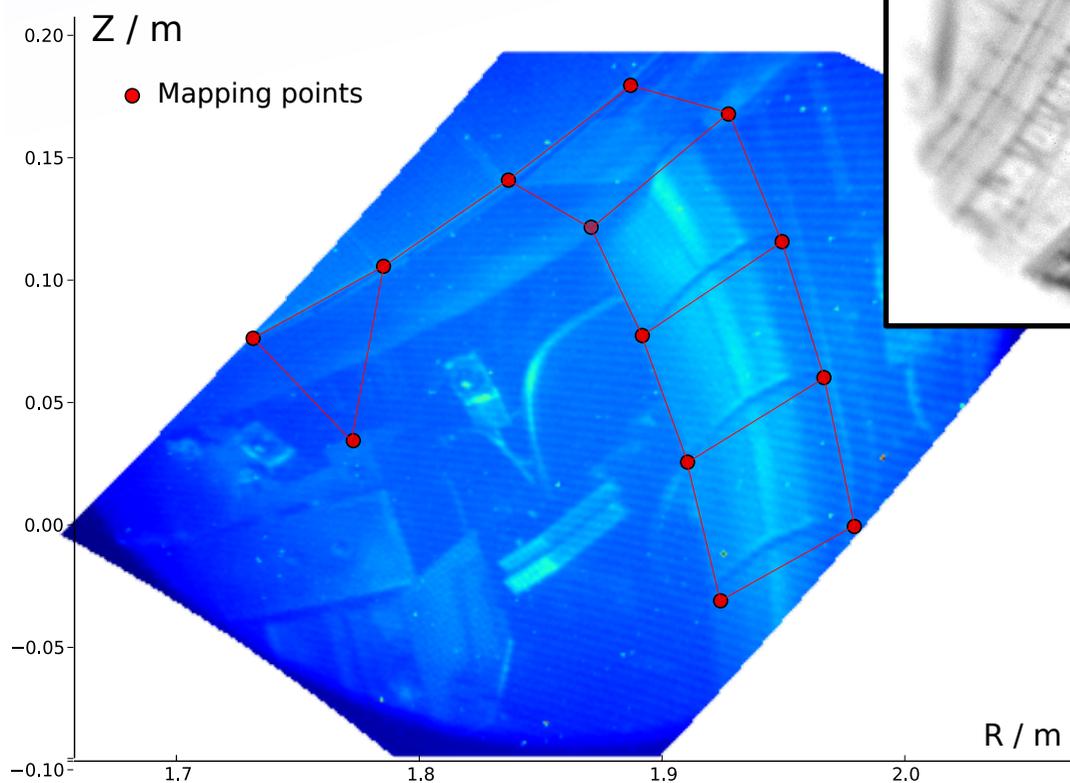
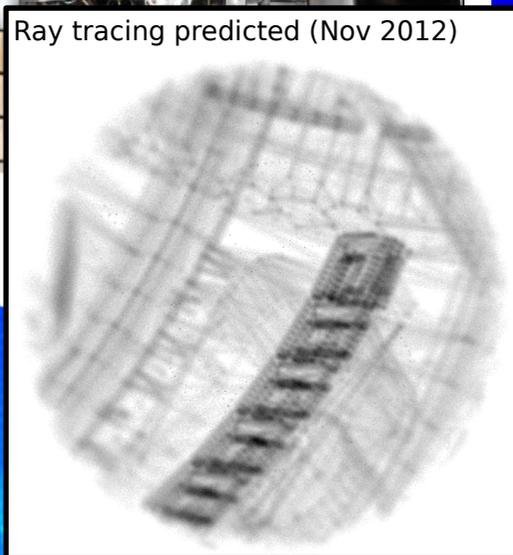
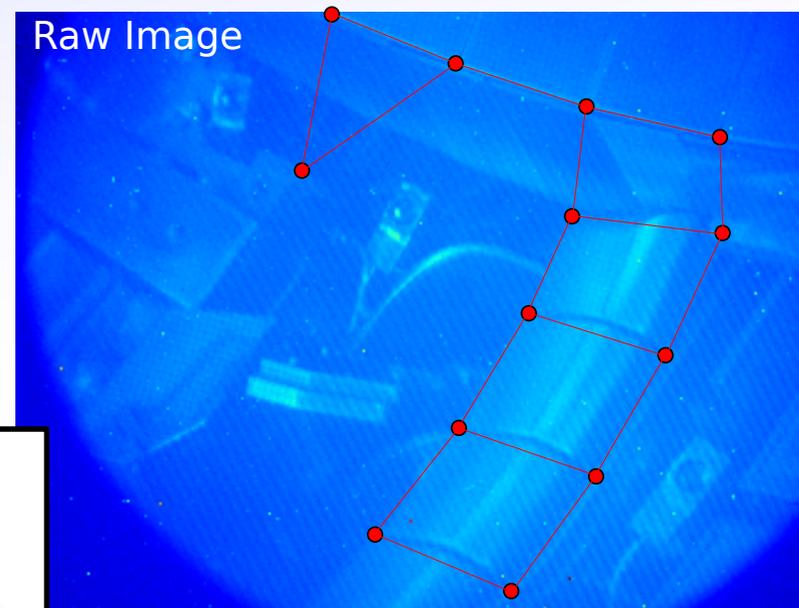
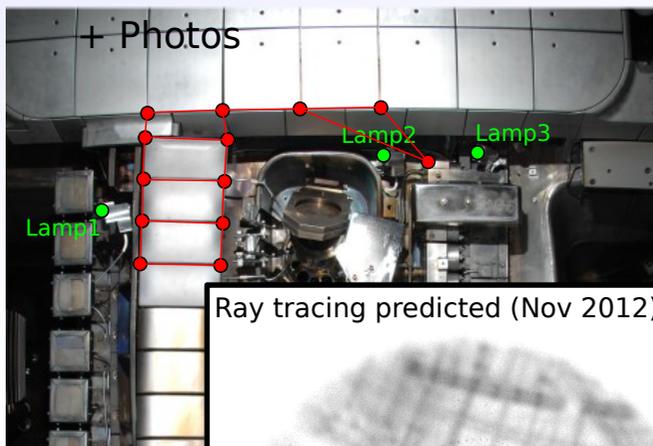
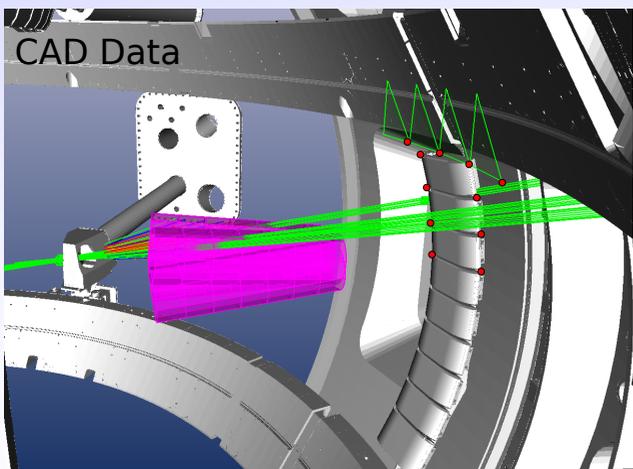
Early results showed expected pitch angle movement during plasma current ramp-up:





# Image Transform

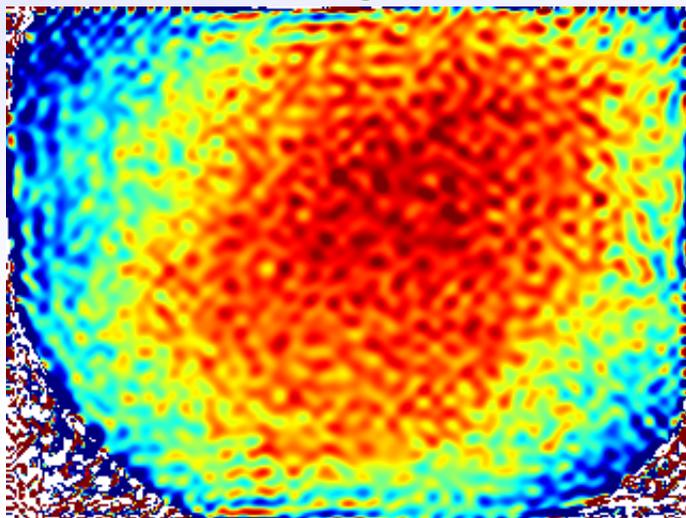
Use background during bright  $D\alpha$  events (e.g. disruptions) for mapping the image to R,Z coordinates at the beam:



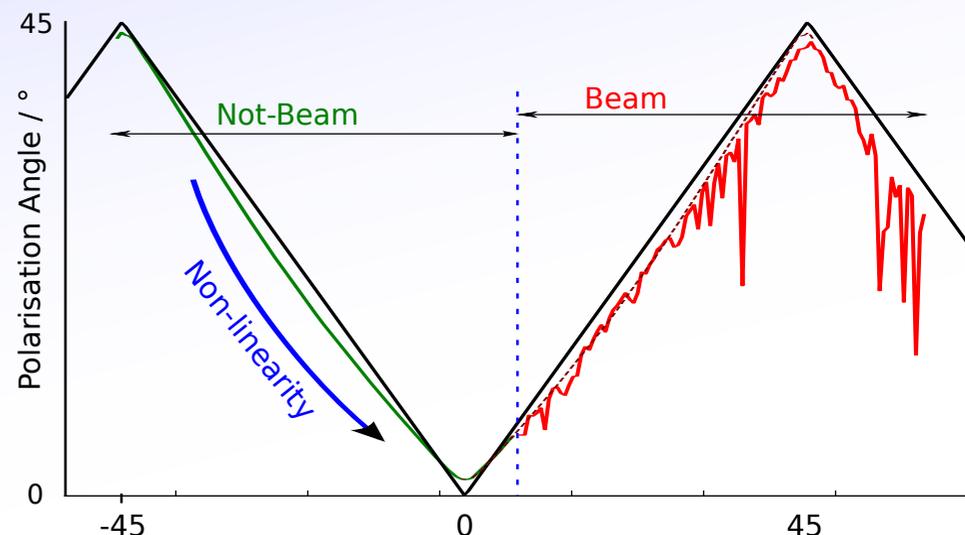
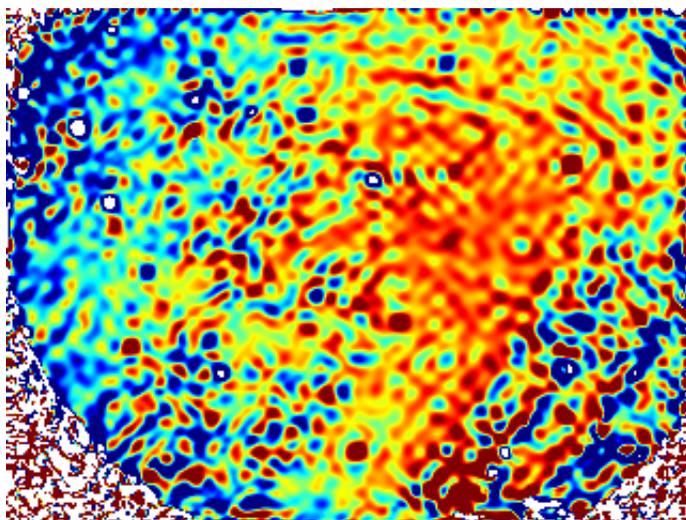
# Intrinsic Contrast Calibration

Calibration of non-linearity requires the polariser to be in place during a plasma, with the beam on:

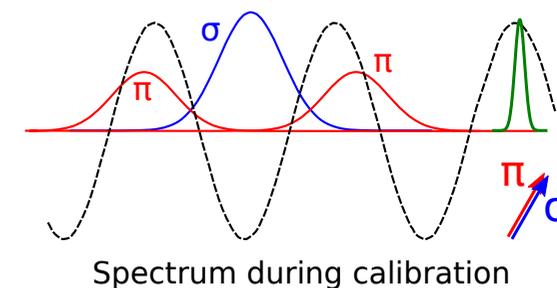
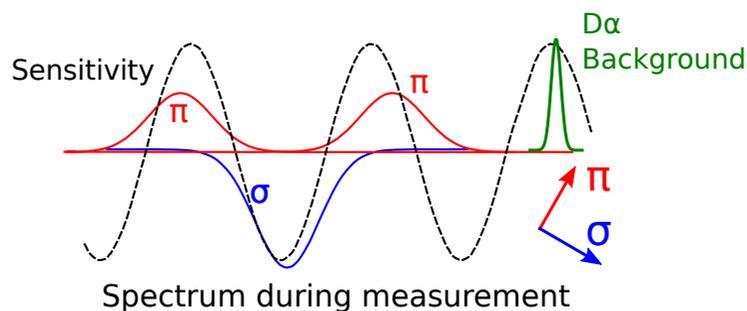
Not-Beam ( $D\alpha$  background)



Beam



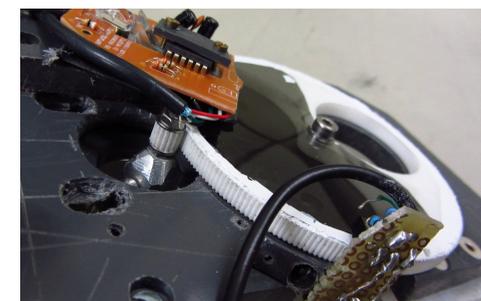
Beam has very low contrast, because the system is optimised for  $\sigma$  and  $\pi$  interferograms to add when polarised at  $90^\circ$ . With calibration polariser, they are parallel so the system is now 'optimised' for them to destructively interfere.



Background  $D\alpha$  is a single spectral peak, so remains strong. The January calibration data is mostly from the background, so not reliable.

Improvements for April:

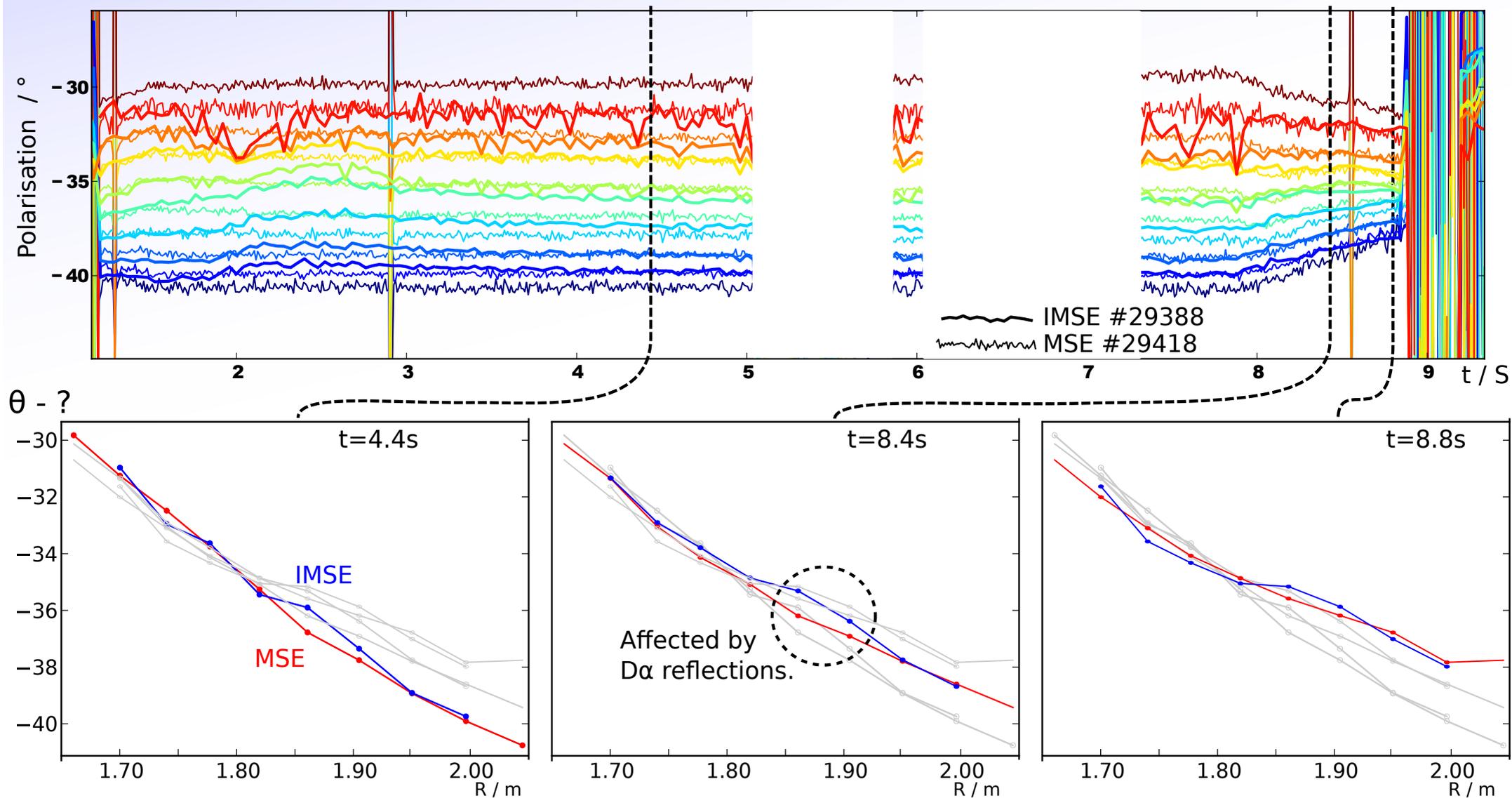
- 3D printed polarisation wheel for better accuracy.
- Increase stepper motor voltage so it works during plasma. (It stopped in Jan due to PF coils fields)
- Calculate best position of polariser for calibration.





# Direct MSE - IMSE comparison

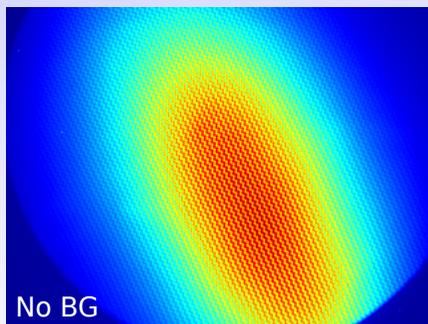
To compare the IMSE/MSE systems directly the same pulse was run the following week with the MSE system. Used the MSE's a polarising wheel as a common zero reference, so the offset should be correct but 1.1° offset is still required:



- ✓ Temporal evolution agreement is very good.
- ✓ Spatial agreement is reasonable (despite calibration problems)
- ✗ An unidentified offset of 1.1° remains.

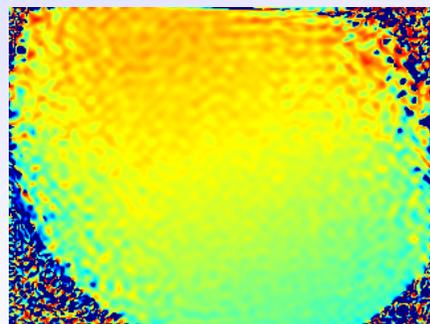
# Reflections - Background D $\alpha$

Image



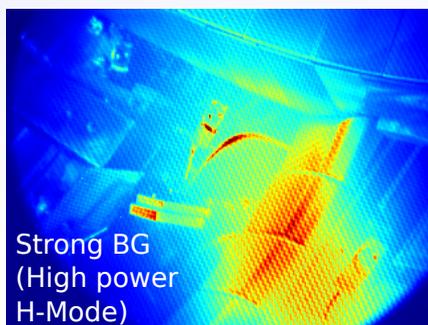
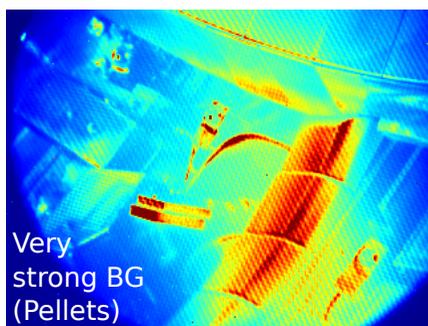
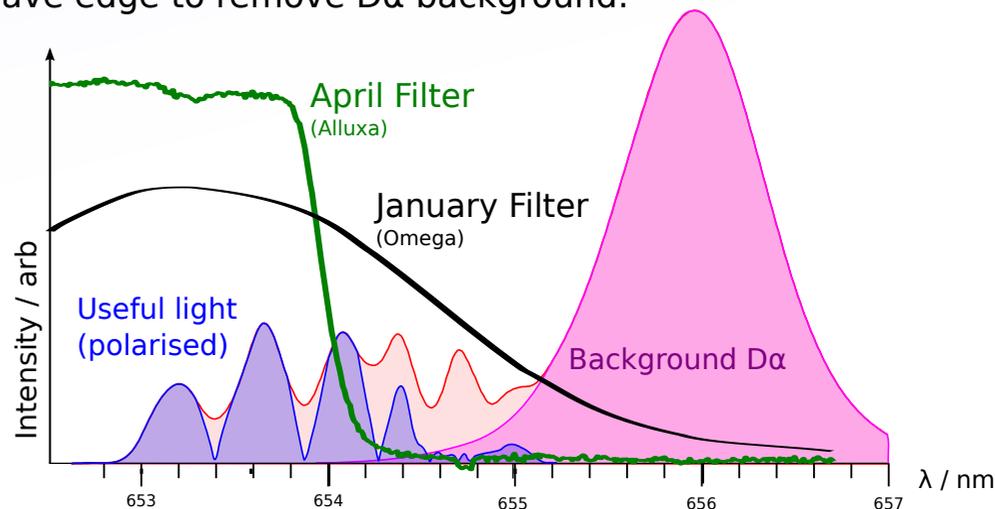
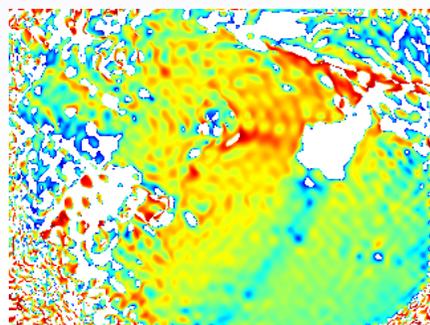
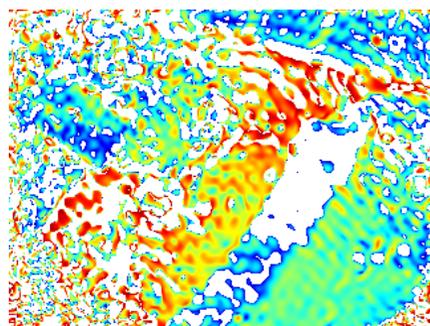
No BG

Polarisation:

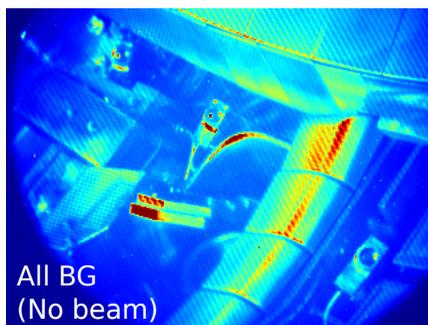
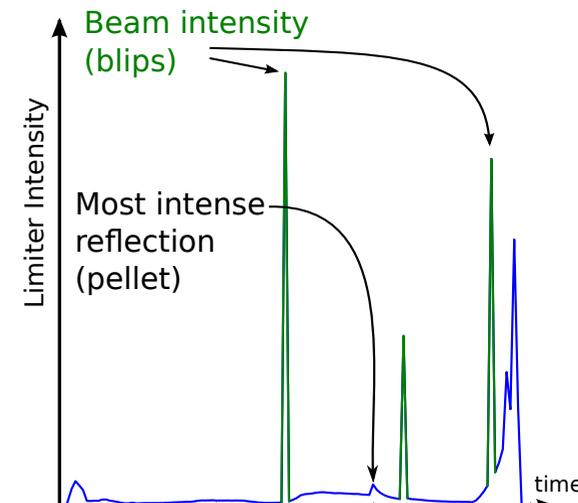
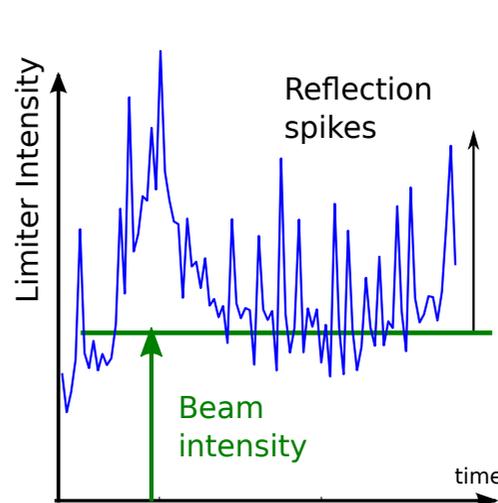
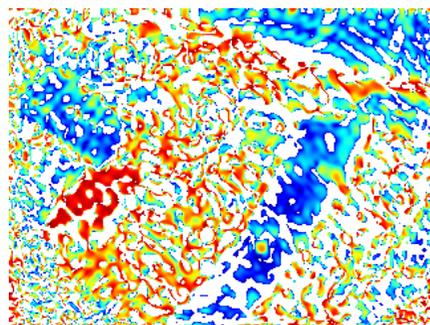


During H-mode / high density operation, D $\alpha$  light reflecting from background structure contaminates the signal.

Particularly bad during detachment and pellets but also makes data unusable during most high power H-modes. New filter purchased for April experiments with very steep short-wave edge to remove D $\alpha$  background:

Strong BG  
(High power H-Mode)Very strong BG  
(Pellets)

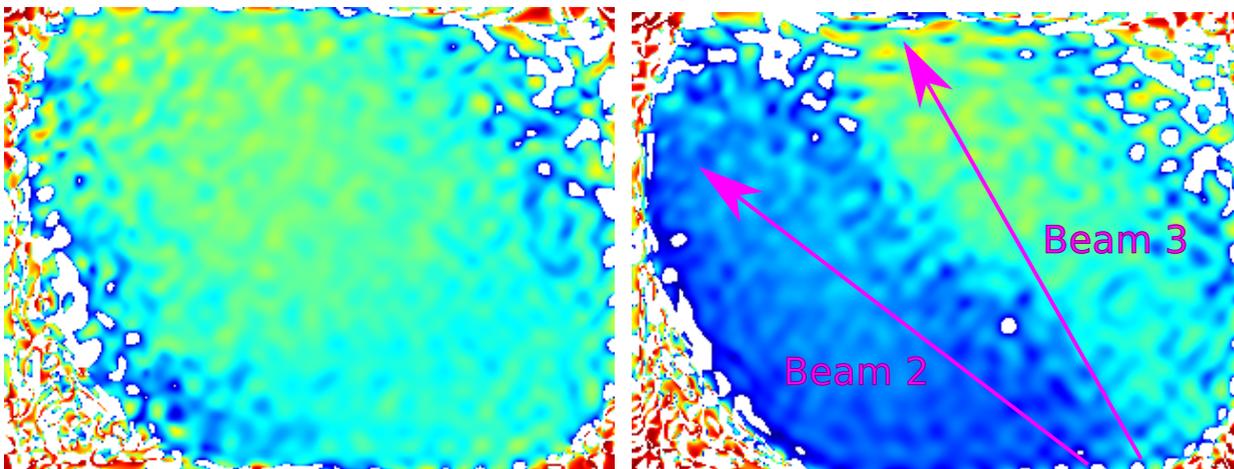
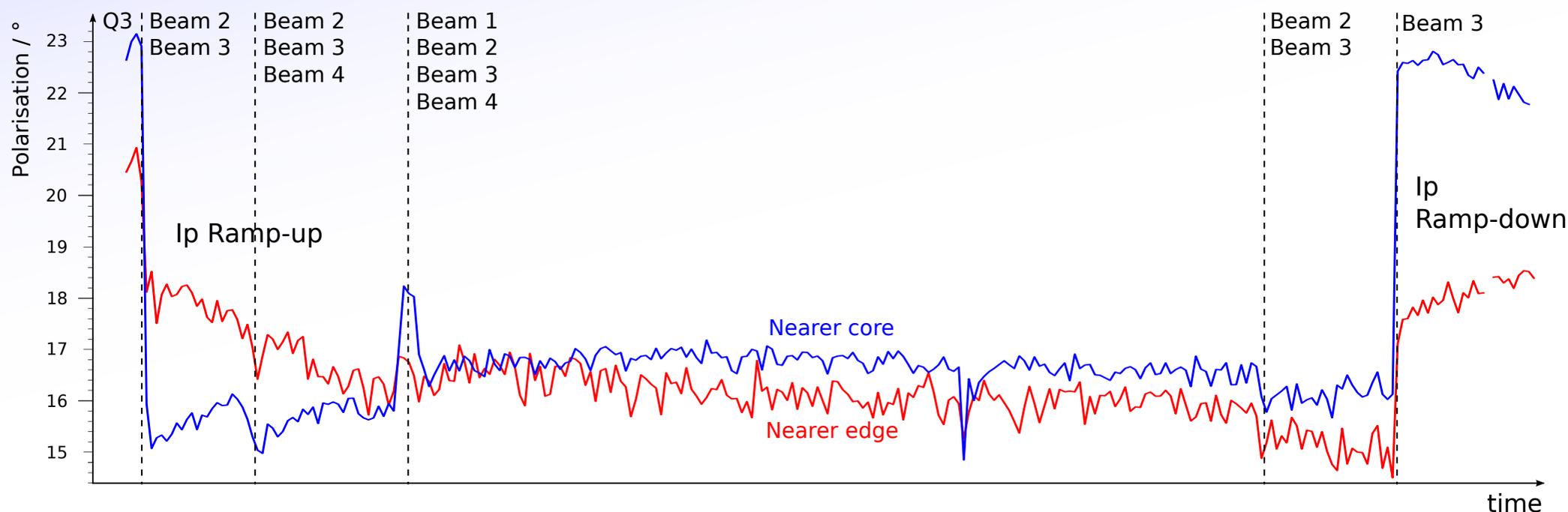
New filter greatly reduced the reflections:

All BG  
(No beam)

# April 2013 Results overview

MSE system does not operate during Hydrogen week at end of campaign (due to change of spectrum) but IMSE is unaffected, so IMSE was reinstalled for this week.

The IMSE worked well, and all modifications were successful. Unfortunately, an accidental machine vent meant most of the week was non-sustained breakdowns. Only a few shots of good data, e.g:



All H-Mode data is unfortunately mixed-beam.

In principle, still possible to use but it will much longer to analyse as it requires knowing the beam geometry and attenuation very accurately.

The beam information is much better separated than my model predicted - so need to improve the beam model and fit the beam waist and divergence.

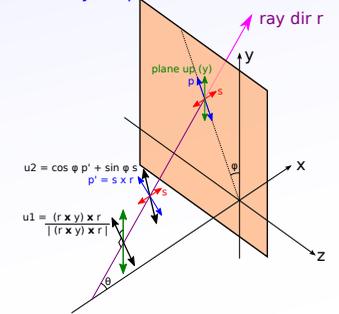


# Absolute Angles

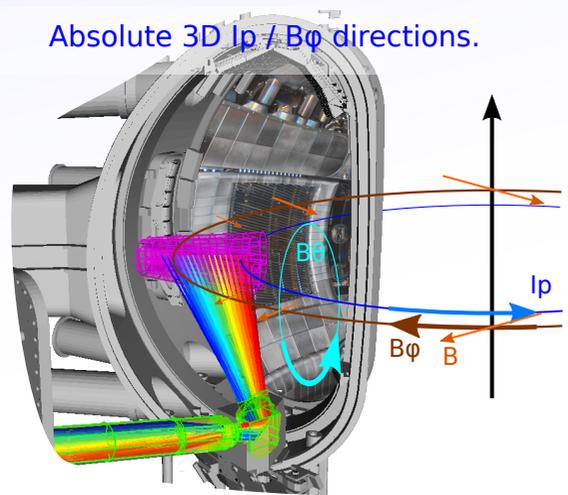
Attempt to predict absolute polarisation angles at IMSE instrument entirely from first principles, to be compared to the measurement later. No calibrations, arbitrary shifts, offsets or flipping of signs to match the observations allowed (proper blind trial)

Will never be accurate enough for analysis but helps to understand the system; identify and document all sign conventions, angle definitions and coordinate systems and to check that there are no unexpected effects. It requires identifying lots of aspects, in absolute terms:

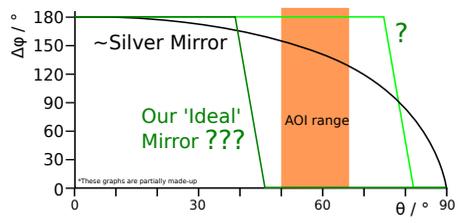
Some subtleties about defining polarisation in 3D (Which way is up?)



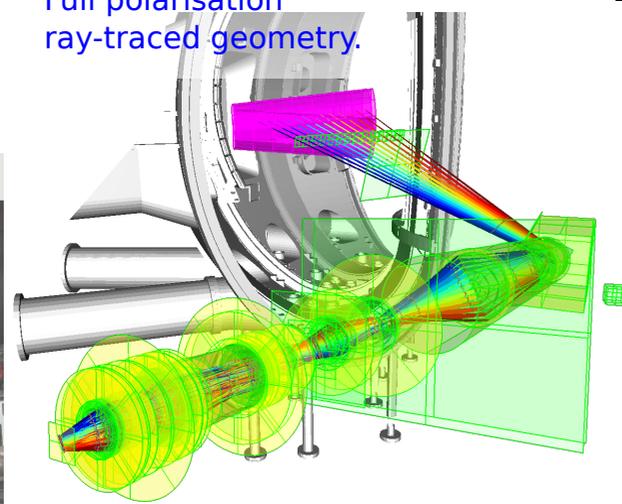
Absolute 3D  $I_p / B_\phi$  directions.



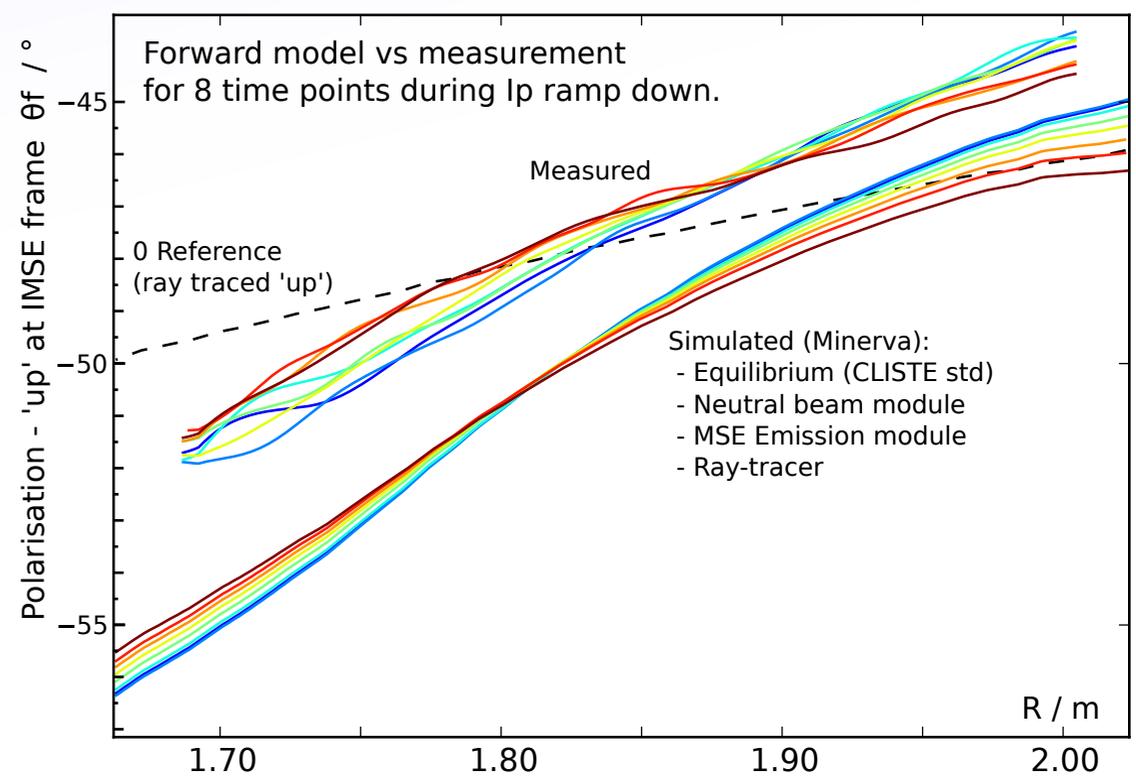
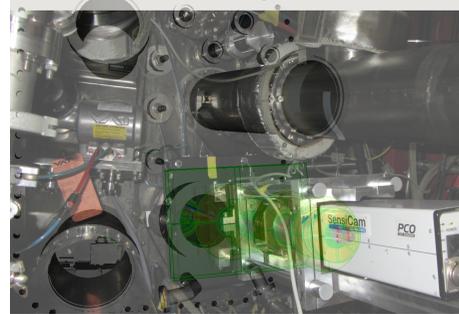
Mirror behaviour (The spherical cyclops in space problem).



Full polarisation ray-traced geometry.



Mounting of IMSE carriage in port flange.



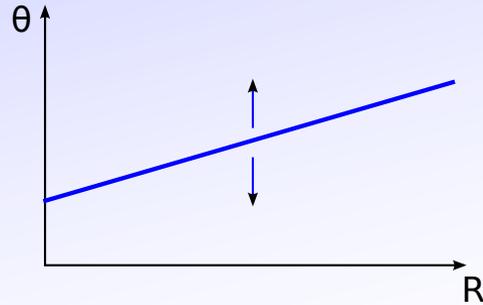
Agreement remarkably good considering inaccuracy of forward model inputs. However, there remains:

- 1.5° variation across field of view (is a problem).
- 2° offset (not a problem)
- Difference in pivot position

(can only be equilibrium or positioning error)  
The optics Faraday effect not yet included ( $\pm 1.4^\circ$  offset)

# Calibration Strategies

### Global offset

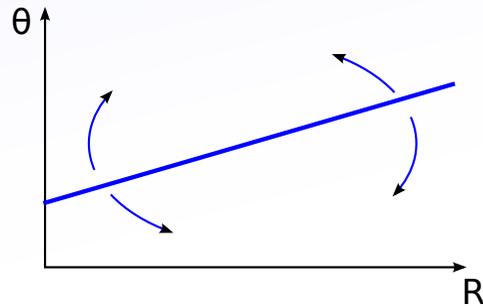


Expected and to some extent acceptable. To my knowledge, no MSE system has ever been successfully absolutely calibrated (without plasma and an equilibrium code).

Strategies:

- Measured background Zeeman-split  $D\alpha$  polarisation.
- Consistency for plasma with different beams (Different  $\mathbf{V} \times \mathbf{B}$  projections).
- Small low  $I_p$  plasmas where the edge (known  $B\theta$ ) can be seen.

### Field of view variation.

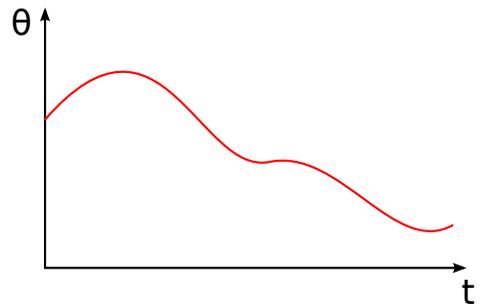


MSE system has an offset for each channel. IMSE should in principle remove this requirement. Remains to be clearly demonstrated.

Strategies:

- Consistency between different plasmas.
- Zeeman split background  $D\alpha$ .

### Long term drifts



Stability over long periods of offset - drifts of rotation in the forward optics (coatings etc) and isolation of the Faraday effect in the forward optics.

Strategies:

- Zeeman split background  $D\alpha$ .
- Limiter reflection polarisation (linearly polarised for some reason).



# Status

- ✓ Modelling for AUG spectrum and view.
- ✓ Plate and optics design for AUG.
- ✓ In-depth performance tests for polarisation measurement.
- ✓ Control, calibration, capture, demodulation and processing software.
- ✓ Built and installed IMSE at ASDEX Upgrade.

- ✓ Improved non-linearity calibration (rebuild the wheel).
- ✓ Purchase and fit  $D\alpha$  blocking filter to remove reflections.
- ✓ Independent test and measurement of Zeeman split  $D\alpha$ .

- Isolate and reproduce field of view polarisation variation.
- Infer absolute offset for Jan and April data.
- Use to infer current profiles via Current Tomography.

- Consider installing permanent system to run alongside MSE.
- Devise method of absolute angle calibration.
- Complete ellipticity recovery investigation (needs better test sources).
- Model Stark-Zeeman coupling (can ellipticity tell us something?)
- Deeper investigation of  $\theta \rightarrow J\phi$  inference methods (including equilibrium codes)

EPS + Satellite diagnostics conference.

- ✓ Modelling and image generation for W7X.
- ✓ Ability to infer polarisation angle for W7X images.
- ✓ Ability to infer parallel current for W7X.
- ✓ Modelling of AUG MSE emission spectrum/polarisation.
- ✓ Modelling of generated images and ability to infer polarisation angle images.
- ✓ Assess ability to infer axisymmetric current from polarisation images.
- ✓ Ray tracing of existing MSE optics to see the image delivered.
- ✓ Assess lens options for optics coupling.
- ✓ Calculate and order required crystals (for middle range of lenses)
- ✓ Calculate and order filter
- ✓ Investigate non-orthogonal fringes idea (works great)
- ✓ Build polarisation test setup (simulated spectrum, OK, getting better filters).
- ✓ Measurement principal test (Zeeman splitting).
- ✓ Software/Methods for exact alignment of plates.
- ✓ Design support structure.

- Model polarisation effects of MSE forward optics.
- Spectrum+Image model for MAST.
- Calculate expected absolute light level (hence max absolute frame rate).
- Find optimum optical setup (fielding etc).
- Interface to objective auto-focus.
- Plasma-based absolute calibration method/check.
- Add Stark/Zeeman coupling to forward model (Ellipticity).
- Reproduce non-linearity in ray tracer (requires full E/O ray splitting to work)
- Improve edge effect / non-periodic FFT demodulation problems.
- W7X assessment based on AUG results.