







Videos and images of fusion plasma properties with coherence imaging spectroscopy and polarimetry.

Technische Universität Berlin, November 2013

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 - 2: Plasma Research Laboratory, Australian National University, Canberra

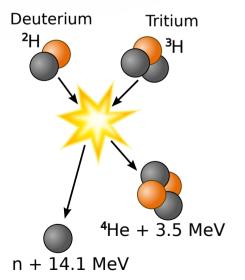


The aim is to produce energy by fusing Deuterium and Tritium nuclei, which produces Helium, a neutron and **lots** of energy.

Why?

- + Clean no radioactive waste products. (Only reactor parts).
- + Carbon free No carbon output from the actual energy production.
- + Abundant Fuel requires only Lithium and sea water, enough for millions of years.
- + Safe Only enough fuel in reactor to sustain reaction we can just turn off the tap.
- + On demand Can turn it up/down as required.
- Very centralised Requires very large, expensive, high-technology machines.

So, it's almost perfect?



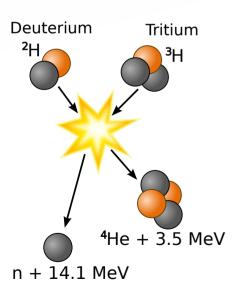


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For sustained thermonuclear fusion, we need enough high-energy collisions of the fuel, that more heat is generated than is lost.

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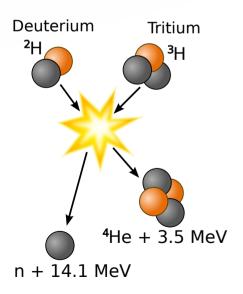


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How can you hold 1×10^8 K without solid materials?



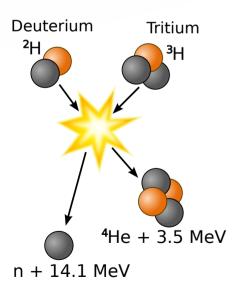


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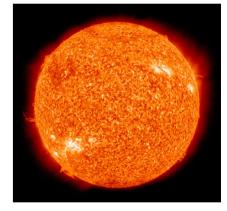
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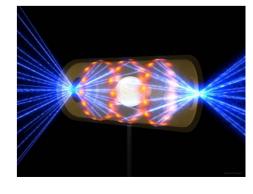
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Gravity?



Inertia?



Magnetic Fields...





At this temperature, the fuels will be a fully ionised plasma, so we can use magnetic fields to confine them.

The electrons and plasma ions move freely along the magnetic field - so the field must be closed and never contact the walls. It must be a torus and the magnetic field must be helical.

Currently two main approaches to do this:



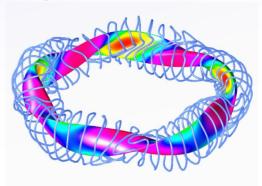
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External coils are complex twisted shapes to twist the plasma.







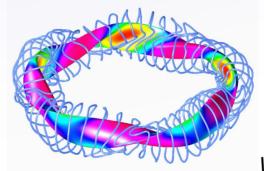
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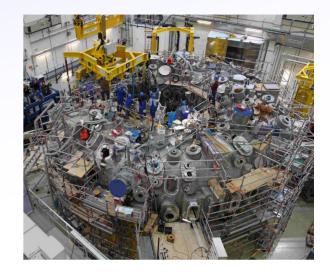
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Wendelstein-7X (IPP Greifswald)





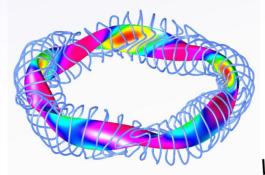
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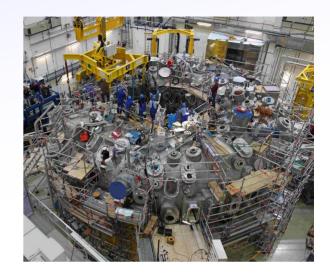
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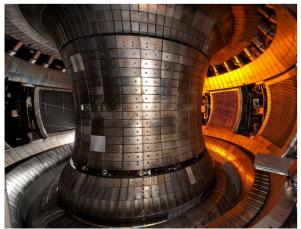
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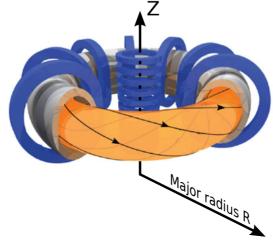


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2) Tokamak:



ASDEX Upgrade (IPP Garching)









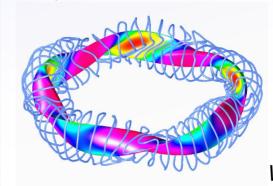
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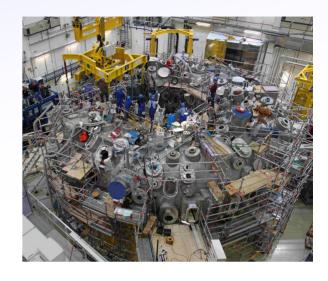
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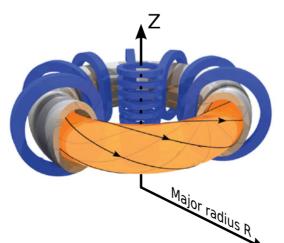
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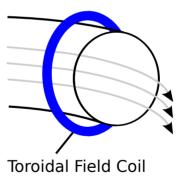
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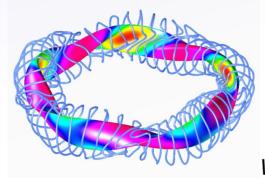
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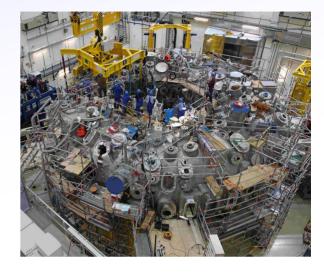
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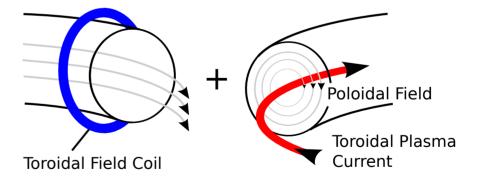
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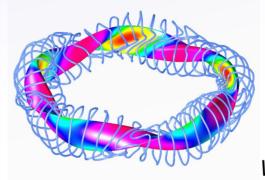
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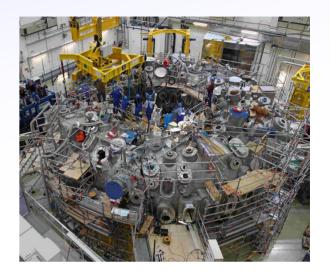
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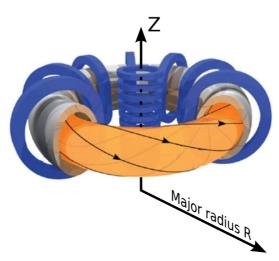
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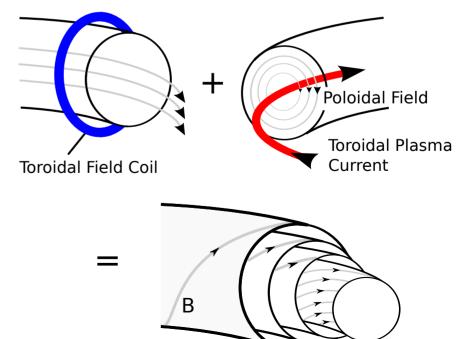
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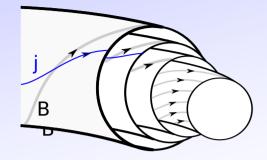
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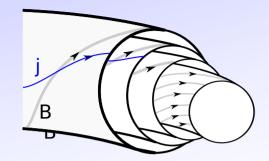




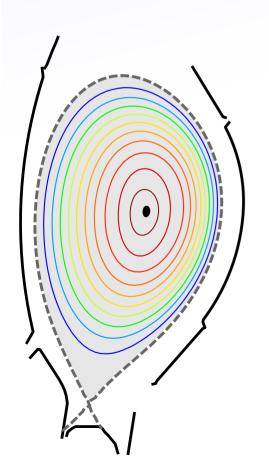
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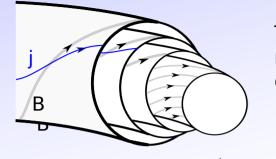


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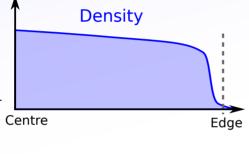
Magnetic Configuration

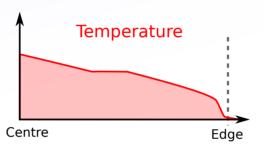


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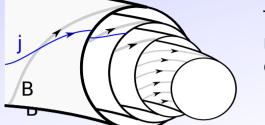




Temperature, density, current* etc are approximately equal around the magnetic surfaces, so knowing the magnetic configuration is vital for interpeting other diagnostics, in order to study the plasma confinement.

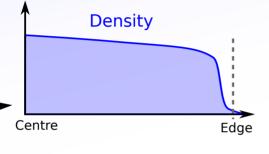


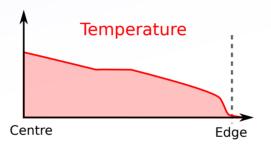


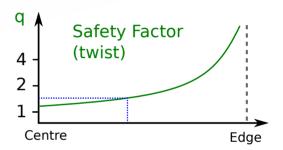


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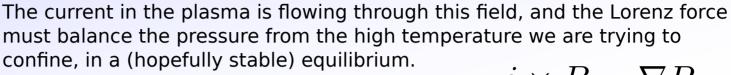
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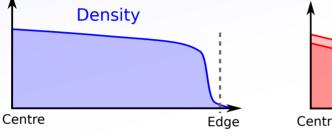
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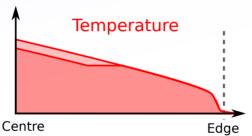


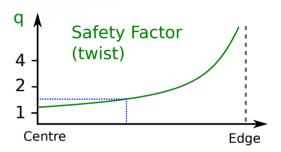
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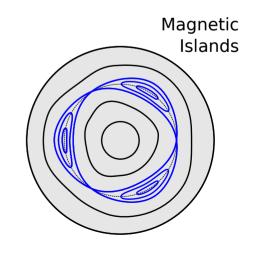






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The twist of the surfaces is so important for the stability, that we call it the 'safety factor' q. Where q is rational, e.g. 3/2, the surface can break up into islands. This increases the transport, so lowers the core temperature, reducing the Fusion reaction rate. Large islands can also disrupt the plasma completely.

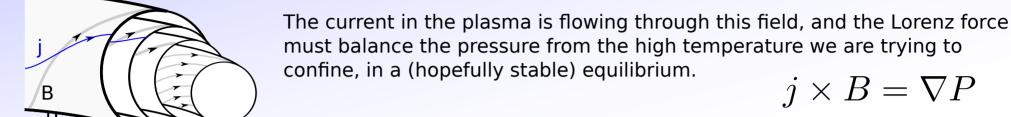


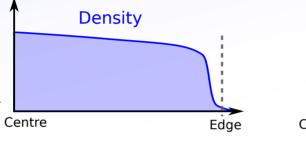


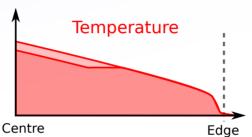
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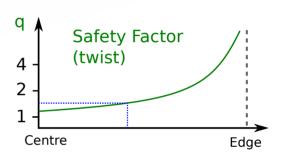


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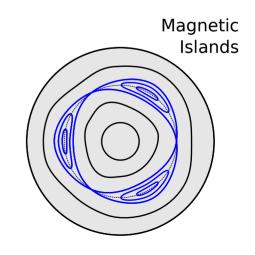




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Recently, control systems have been developed that can drive current to remove the islands, but this requires knowing **q** accurately, in real-time.





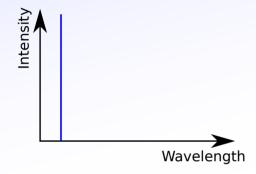


To measure the hot plasma core, we have to examine the emitted radiation and/or particles and use the plasma physics we learn to infer quantities of interest. For example:



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Doppler Spectroscopy: Observe atomic line emission from neutral hydrogen, impurities or laser light scattered by plasma particles.

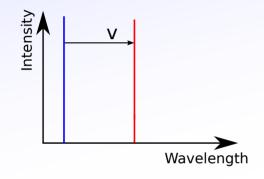






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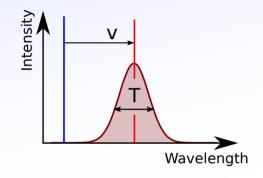
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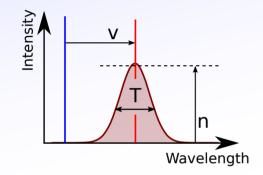


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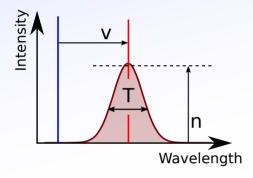


Intensity --> Particle density
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Polarimetry:

Spectral lines are split and polarised by E/M fields:

Zeeman Effect: Magnetic field.

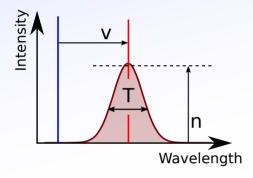
Stark Effect: Electric field.





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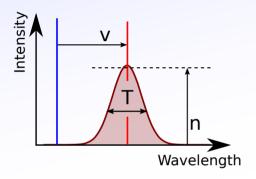
Motional Stark Effect (MSE): Stark effect from Lorentz transformed $\mathbf{E} = \mathbf{v} \times \mathbf{B}$ for fast injected neutrals.





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ASDEX Upgrade Vacuum Vessel

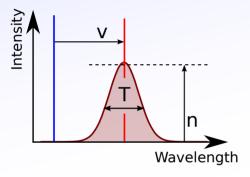




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Magnetic Surfaces Plasma Edge



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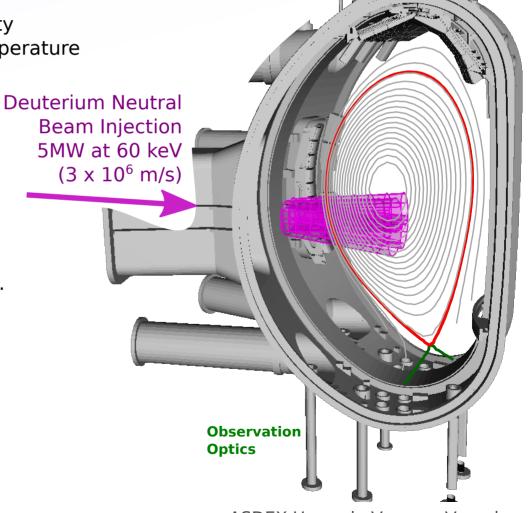
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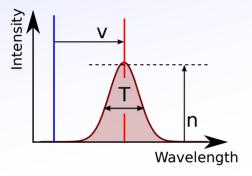


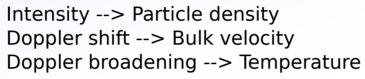


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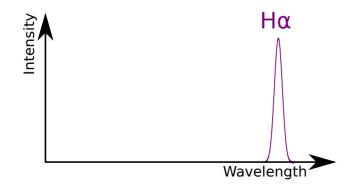
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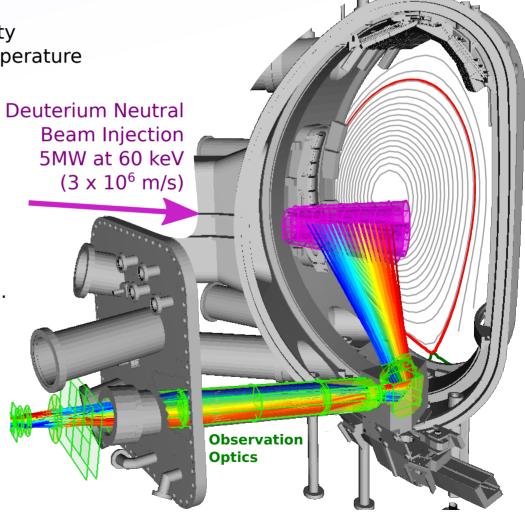
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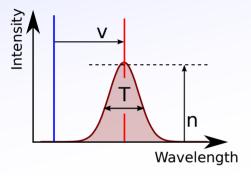


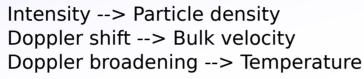


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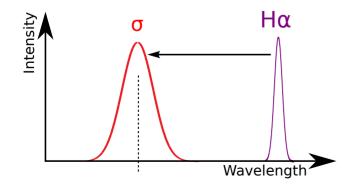
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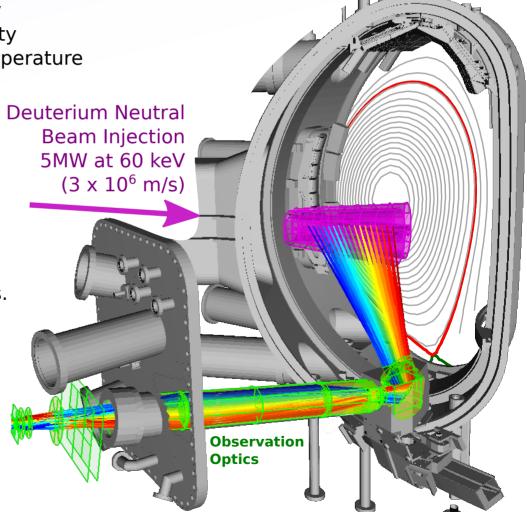
Spectral lines are split and polarised by E/M fields:

Zeeman Effect: Magnetic field.

Stark Effect: Electric field.

Motional Stark Effect (MSE): Stark effect from Lorentz





ASDEX Upgrade Vacuum Vessel

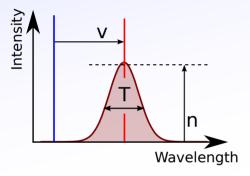


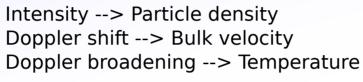


To measure the hot plasma core, we have to examine the emitted radiation and/or particles and use the plasma physics we learn to infer quantities of interest. For example:

Doppler Spectroscopy: Observe atomic line emission from neutral hydrogen, impurities or laser light scattered by plasma particles.

Magnetic Surfaces Plasma Edge





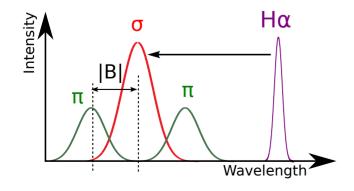
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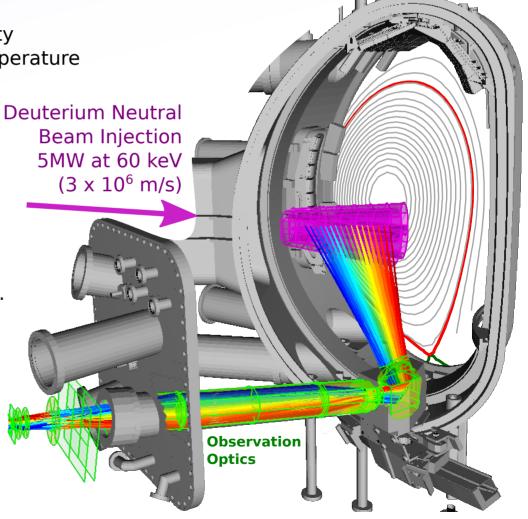
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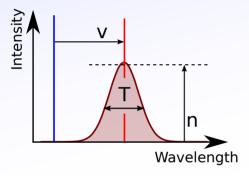


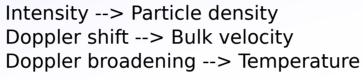


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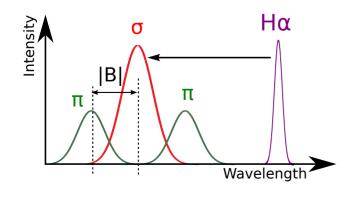
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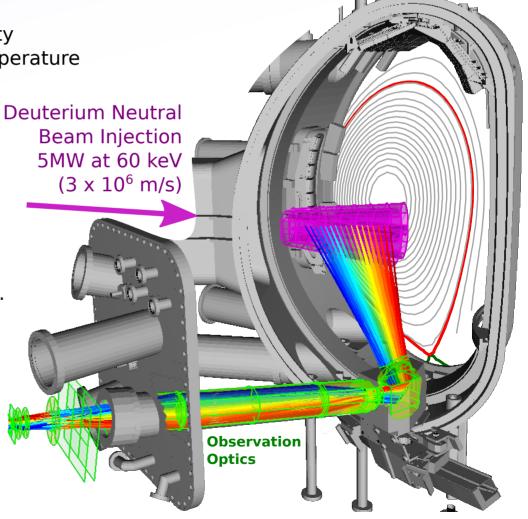
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Polarisation: θ



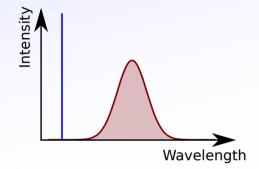
ASDEX Upgrade Vacuum Vessel



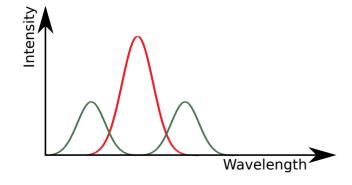
Traditional Systems

Existing systems typically have lots of complex hardware per spatial point:

Doppler Spectroscopy:



Polarimetry:



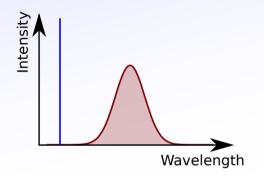




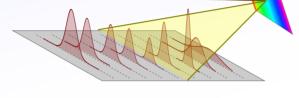
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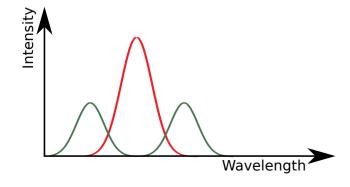


Diffraction grating and CCD camera, or individual detectors (PMTs/APDs)



Low light levels. 1D set of points.

Polarimetry:



or

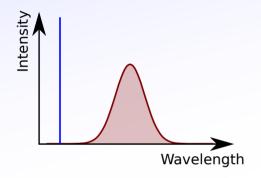




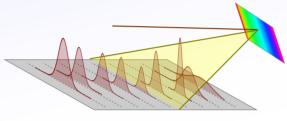
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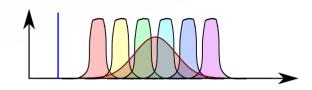


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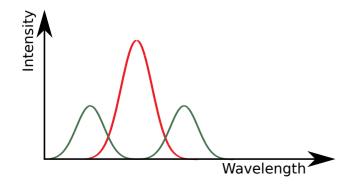
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Individual spectral filters and fast sensitive detectors.



Very complex setup per channel. Low spectral resolution.





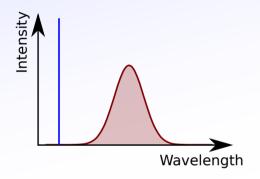
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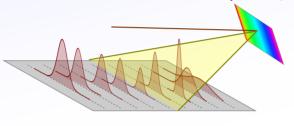
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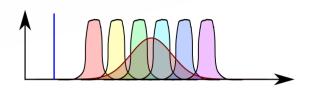


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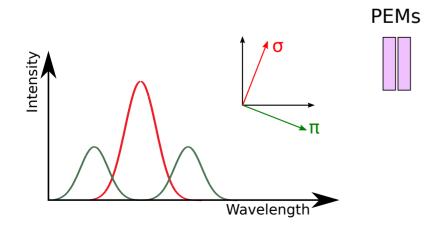
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Motional Stark Effect is usually done with a Photo-elastic modulator (PEM - an acoustically excited crystal)



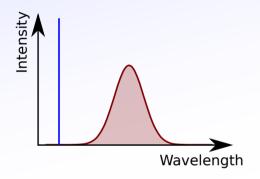




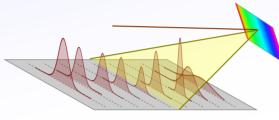
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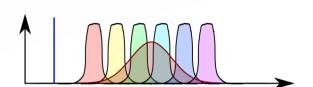


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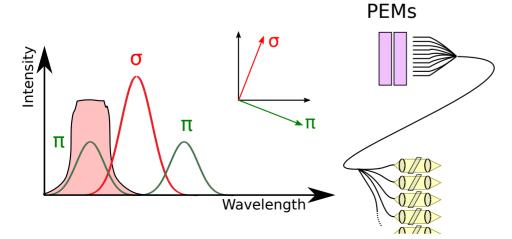
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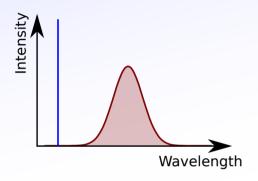




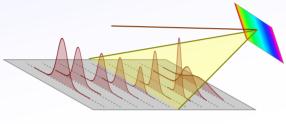
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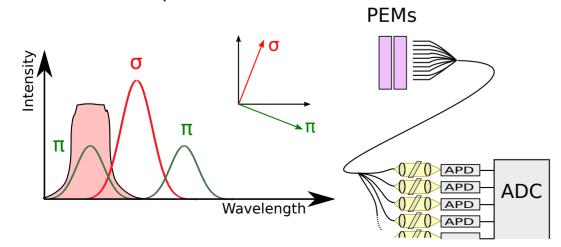
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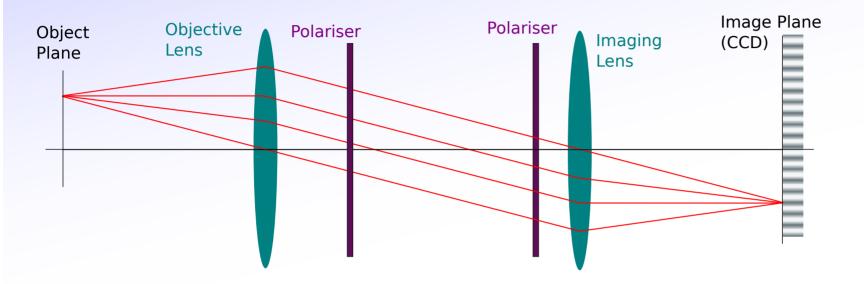
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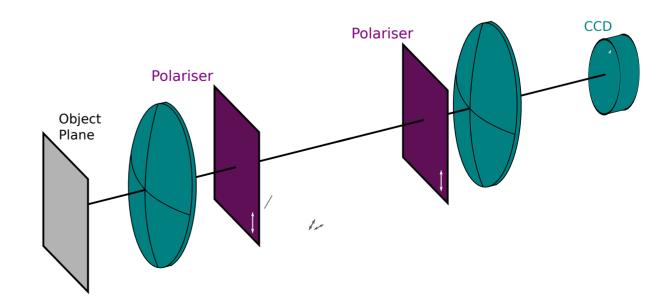




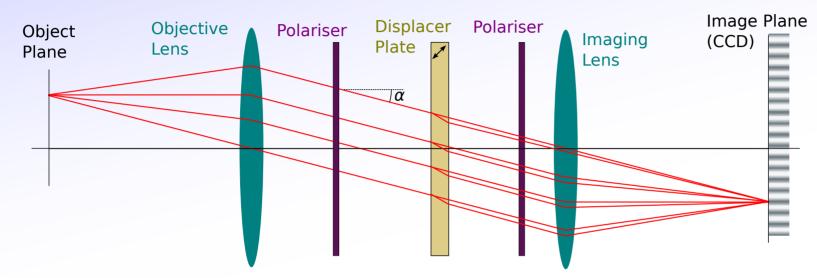


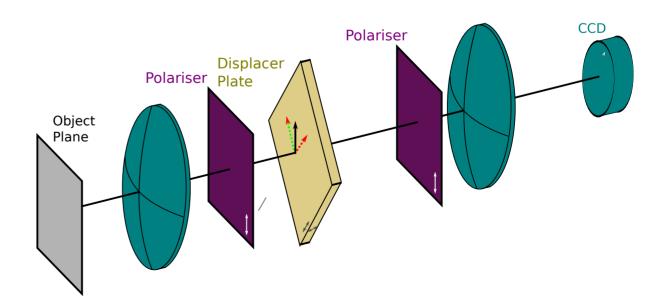




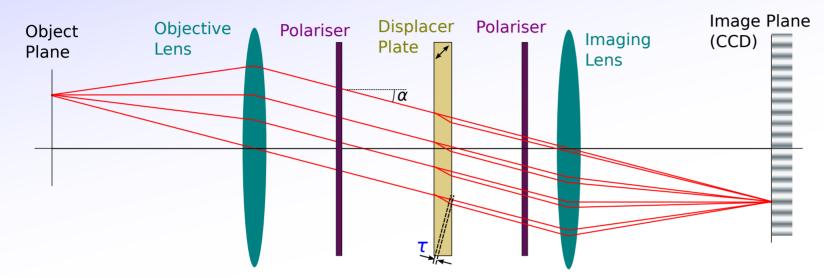


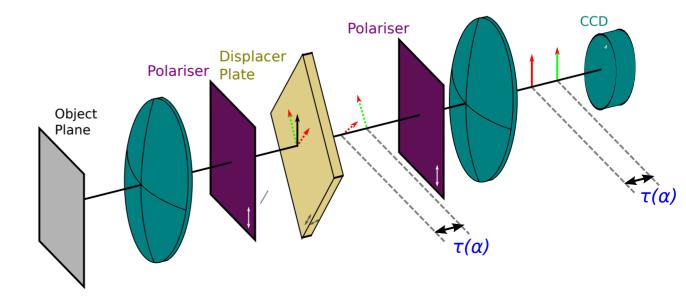






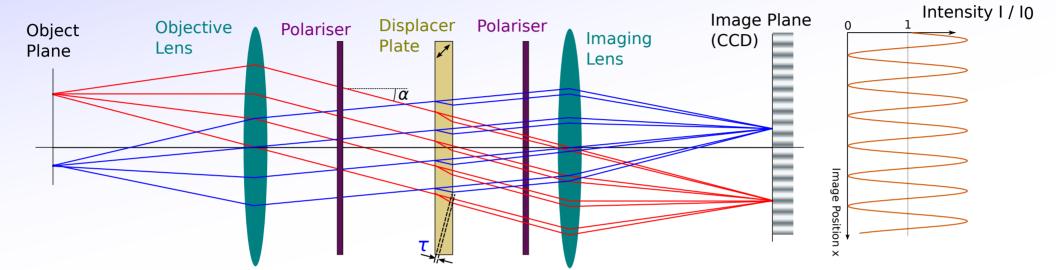


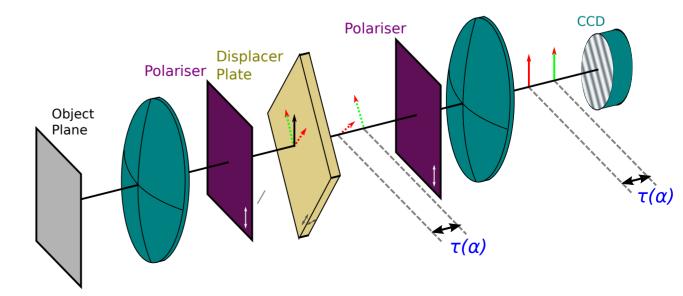






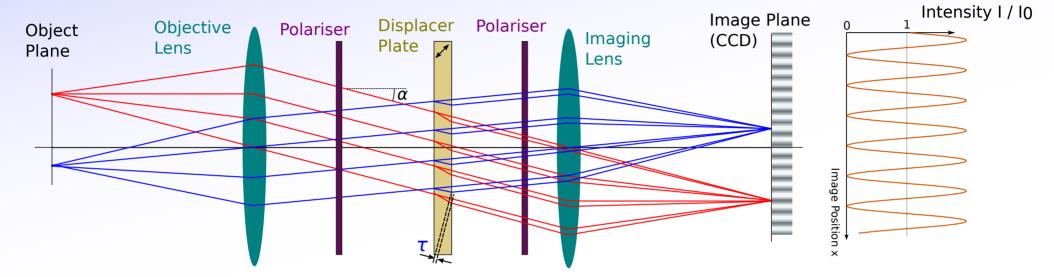


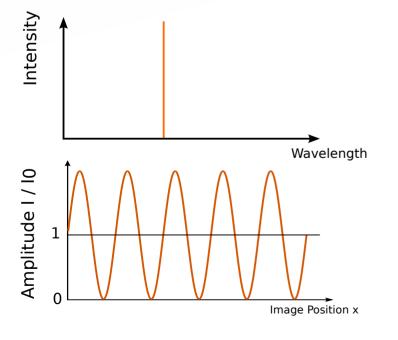


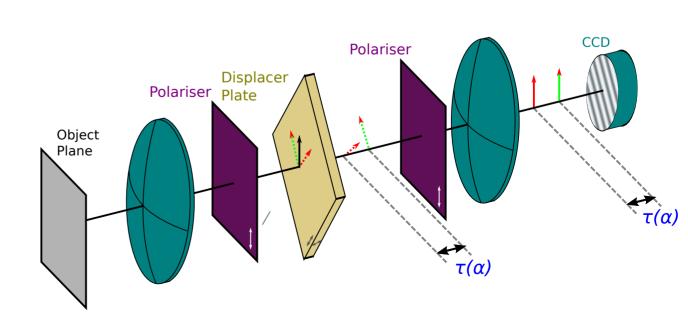




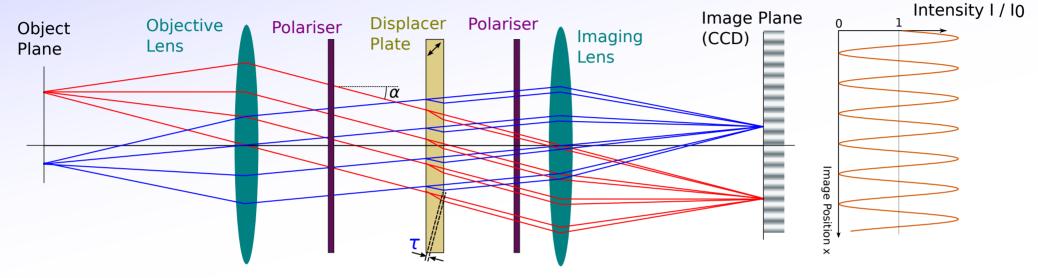


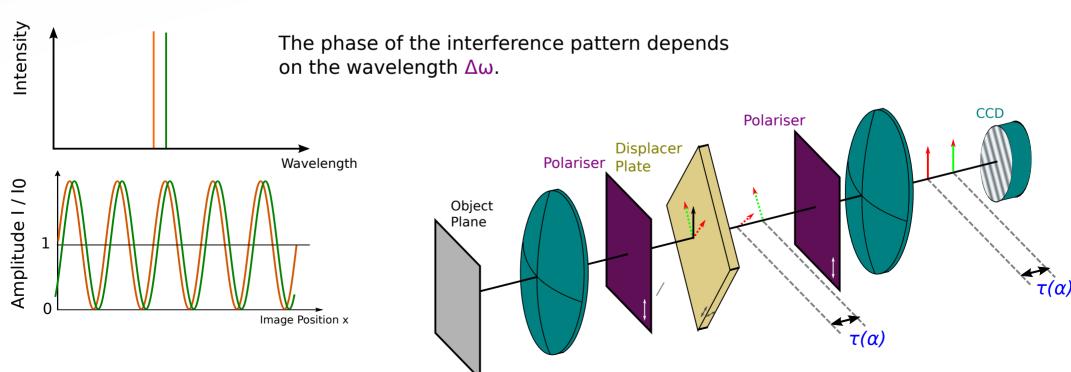




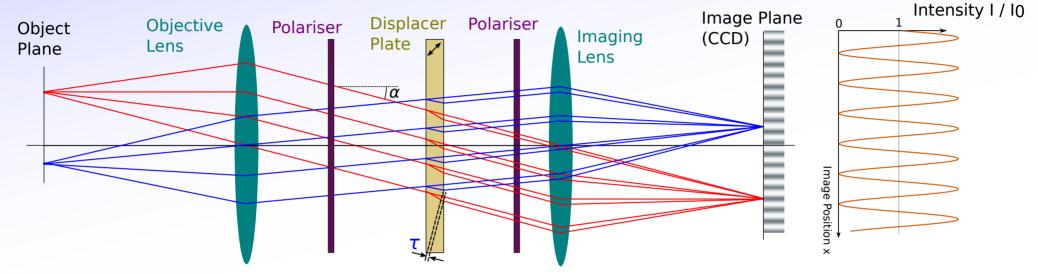


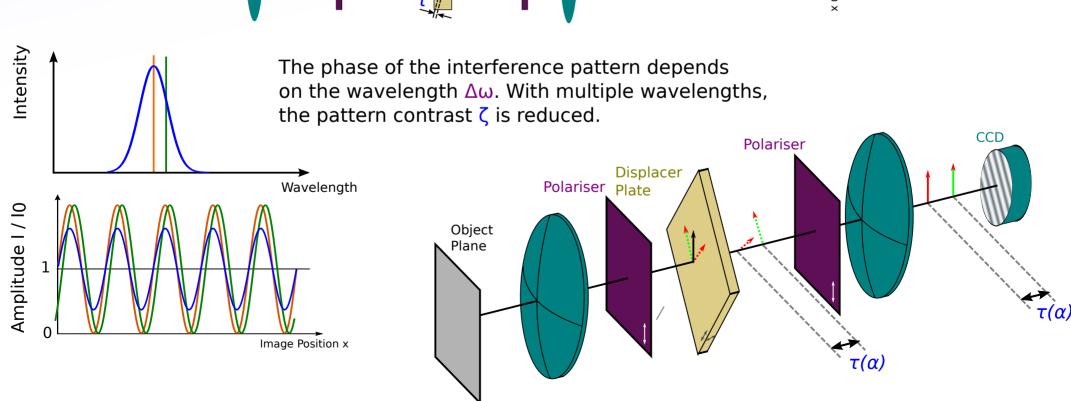






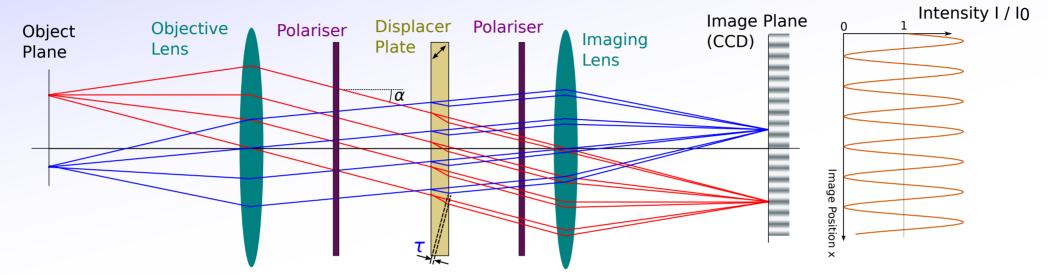


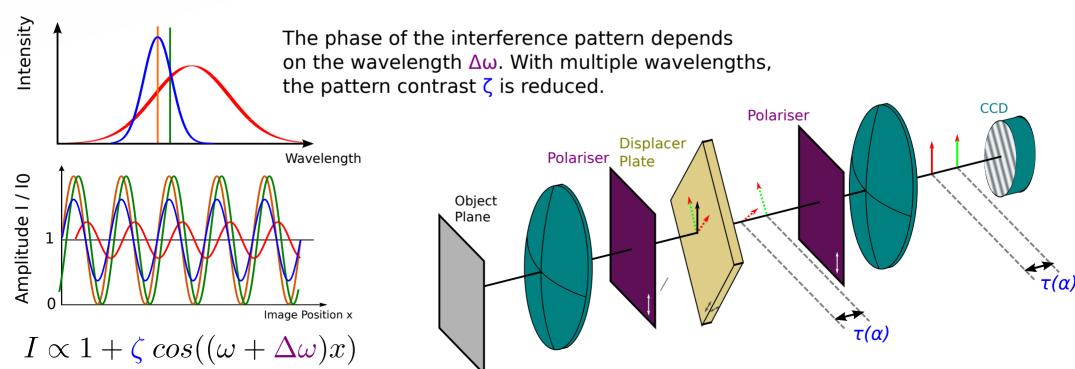














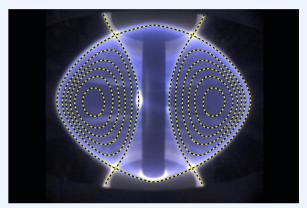


Doppler Coherence Flow Imaging - MAST

Some results of neutral Helium flow in the (relatively) cold edge of MAST:

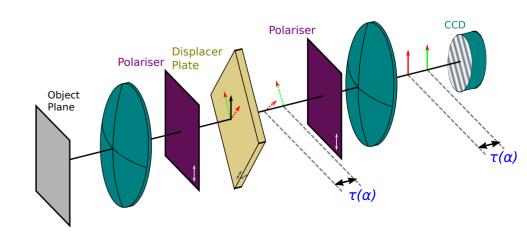
MAST

Mega Amp Spherical Tokamak, CCFE, Culham, UK



MAST is a 'sphereical' Tokamak. The torus has a very small major radius compared to it's minor radius, but is still a Tokamak.

*With thanks to Scott Silburn, Durham University / CCFE [S. Silburn et. al. 40th EPS Conf. on plasma phys. 2013]





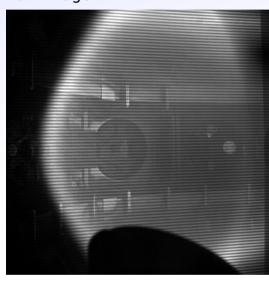




Doppler Coherence Flow Imaging - MAST

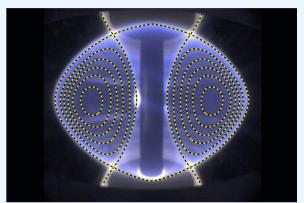
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Raw Image:



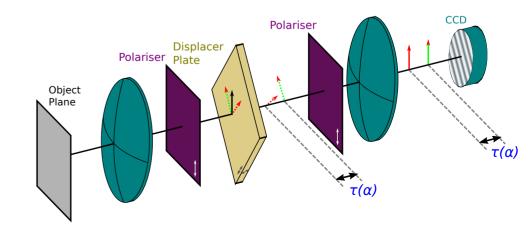
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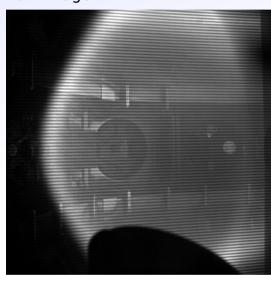




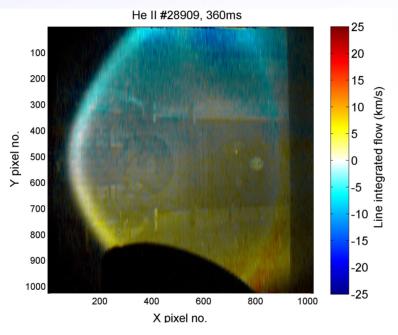
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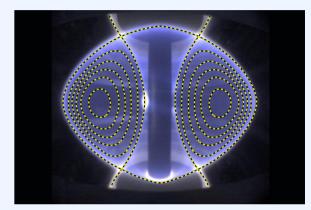


Helium Flow Velocity:



MAST

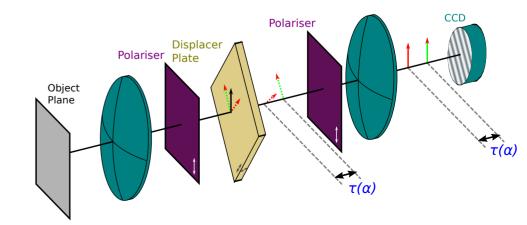
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And some videos...

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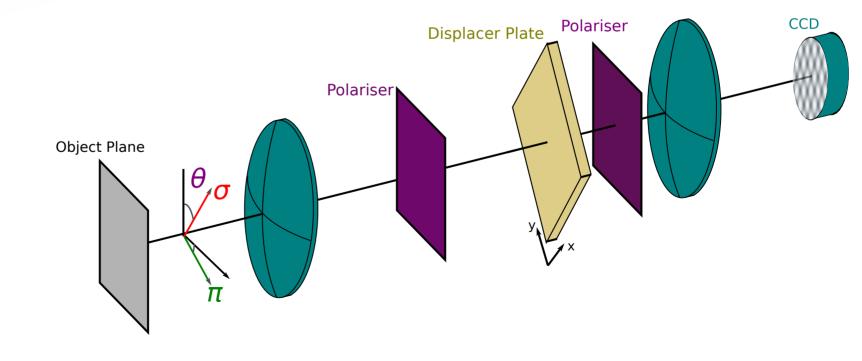






Removing the first polariser gives a dependence on the initial polarisation:





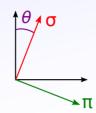


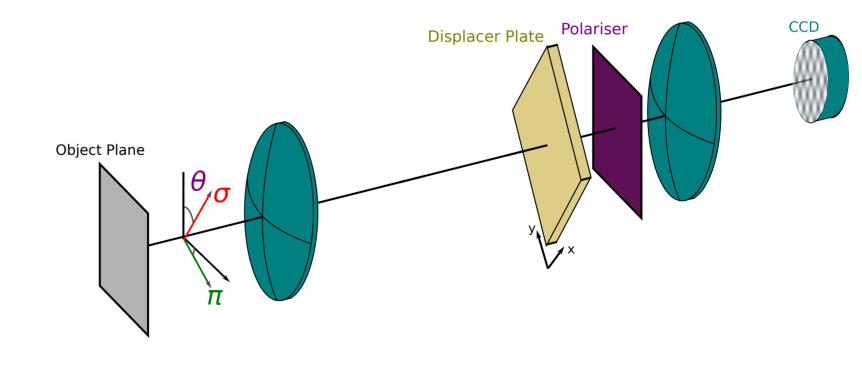


Removing the first polariser gives a dependence on the initial polarisation:

$$I \propto 1 + \zeta \cos 2\theta \cos(x)$$

For the Stark/Zeeman spectrum, the π component is at 90° to σ , introducing a 180° phase shift, so they would cancel.





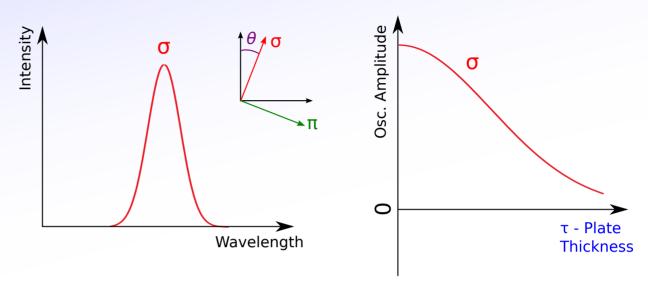


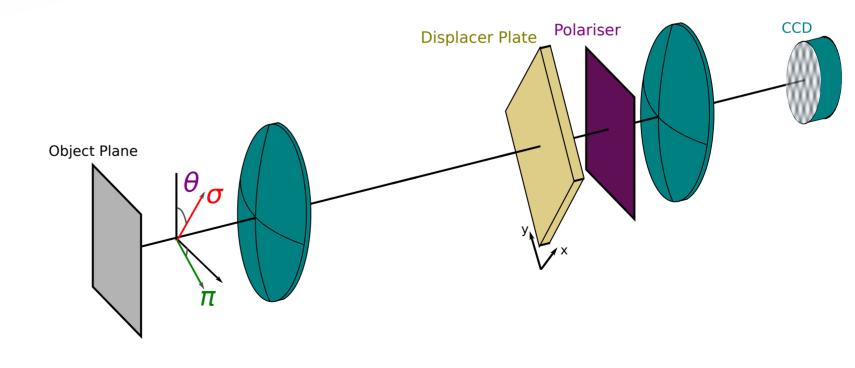


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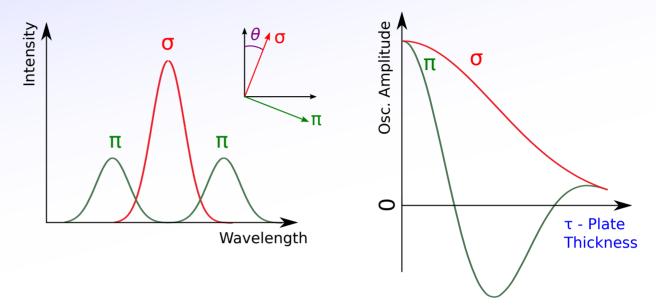


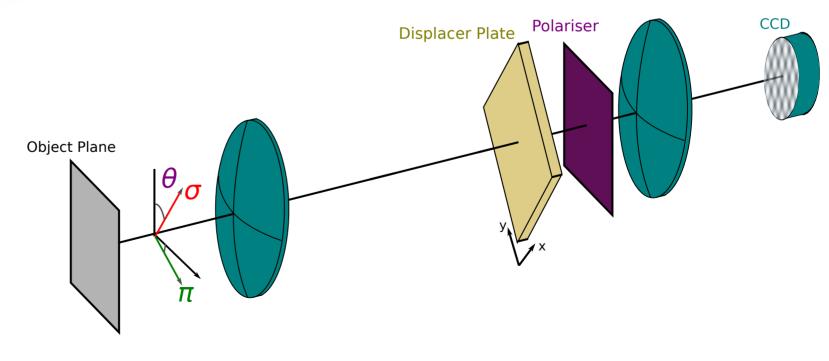
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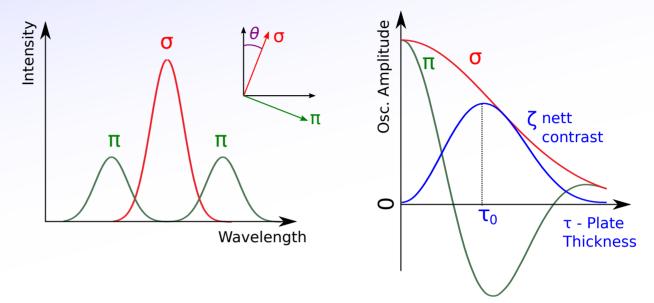


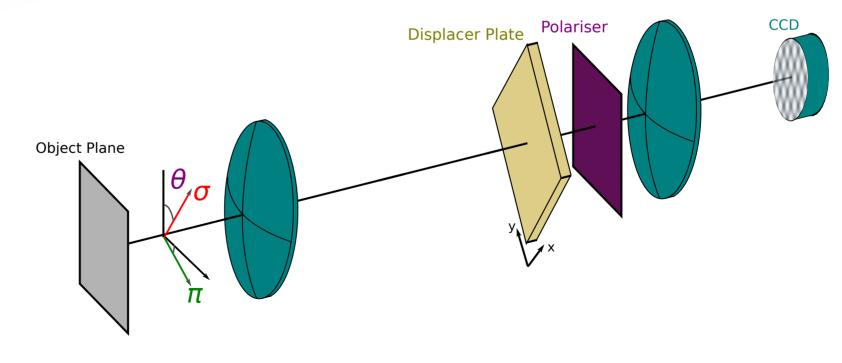
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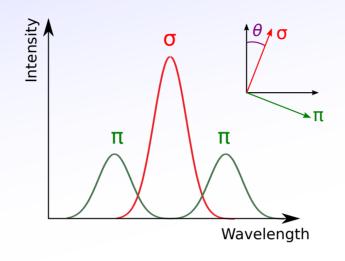
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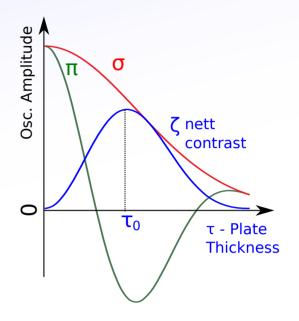
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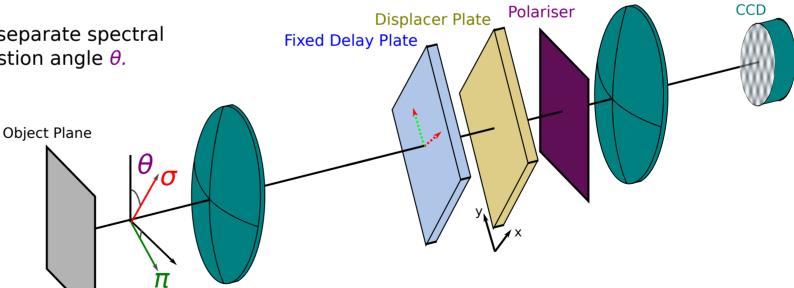
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However, we now need to separate spectral contrast ζ from the polariastion angle θ .











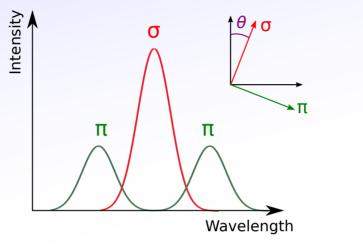
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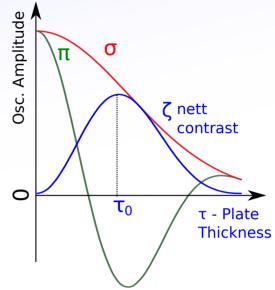
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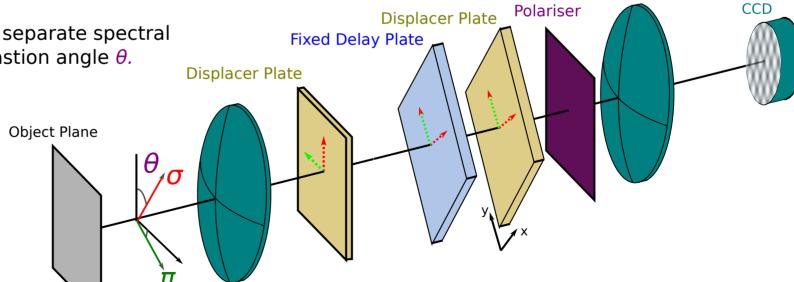
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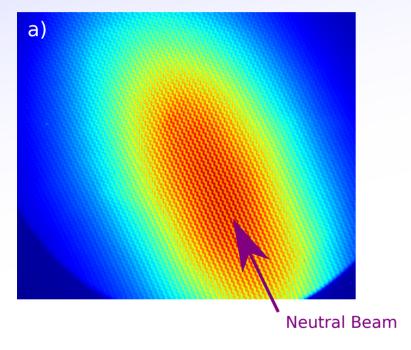


add another displacer at 45°. Combined effect adds 2 extra terms:

$$I \propto 1 + \zeta \cos 2\theta \cos(x) + \zeta \sin 2\theta \cos(x - y) - \zeta \sin 2\theta \cos(x + y)$$

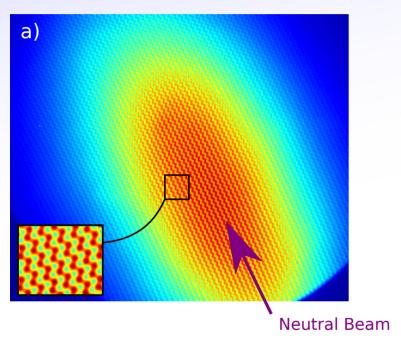


$$I \propto 1 + \zeta \cos(2\theta) \cos(x) + \zeta \sin(2\theta) \cos(x+y) + \zeta \sin(2\theta) \cos(x-y)$$



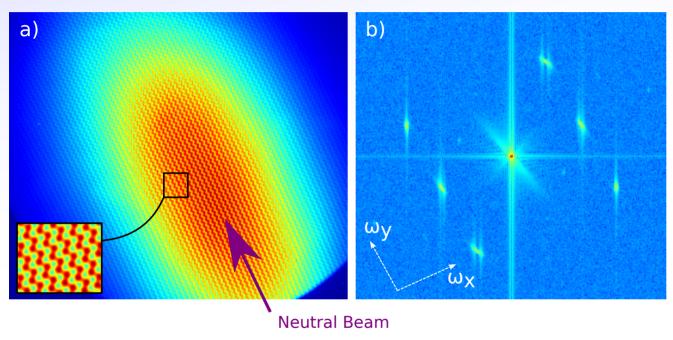


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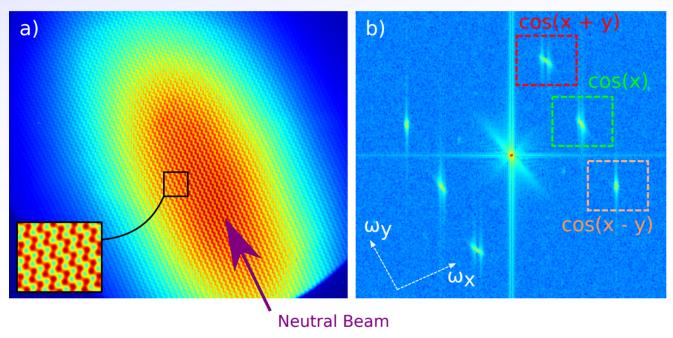


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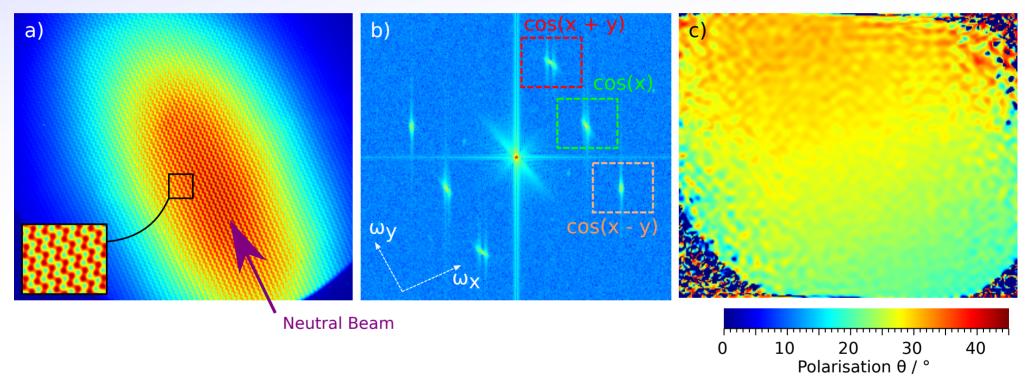


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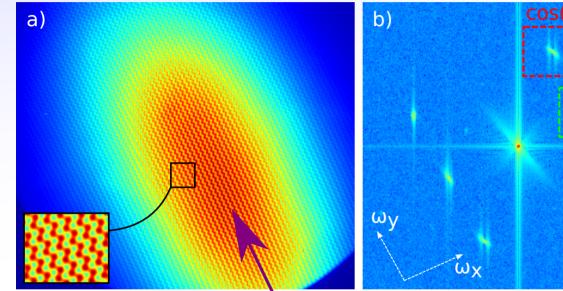
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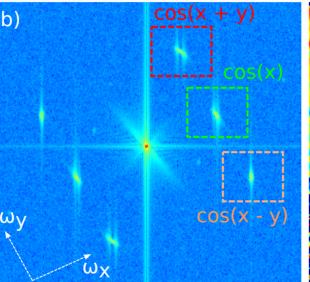


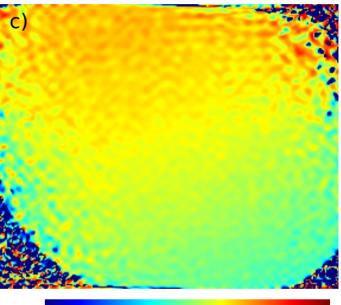


The two orthogonal interference patterns give 3 components in the Fourier transform. We can filter these from the FT and extract the polarisation angle θ :

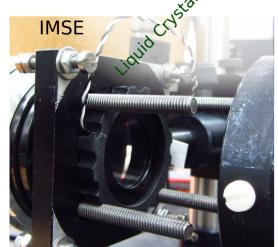
$$I \propto 1 + \zeta \cos(2\theta) \cos(x) + \zeta \sin(2\theta) \cos(x+y) + \zeta \sin(2\theta) \cos(x-y)$$



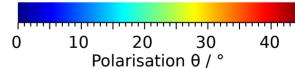








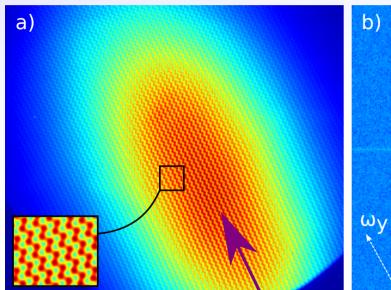
Neutral Beam

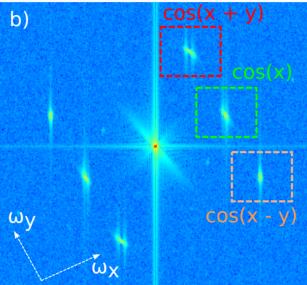


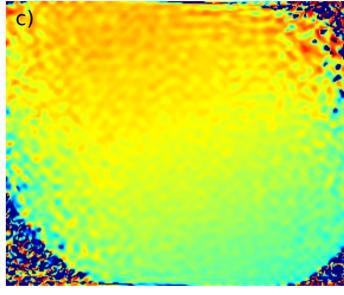


The two orthogonal interference patterns give 3 components in the Fourier transform. We can filter these from the FT and extract the polarisation angle θ :

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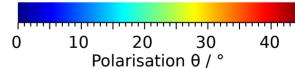




The hardware:



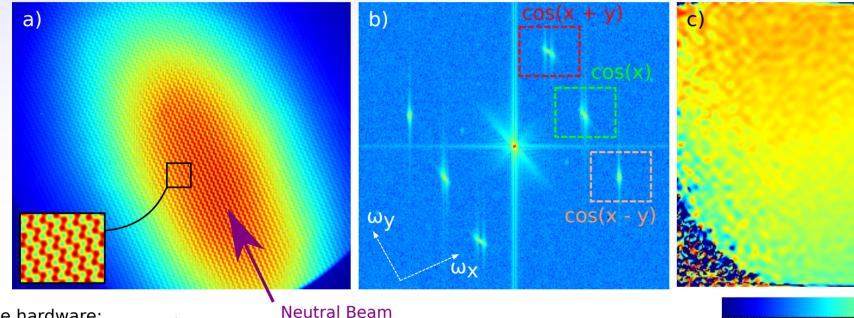
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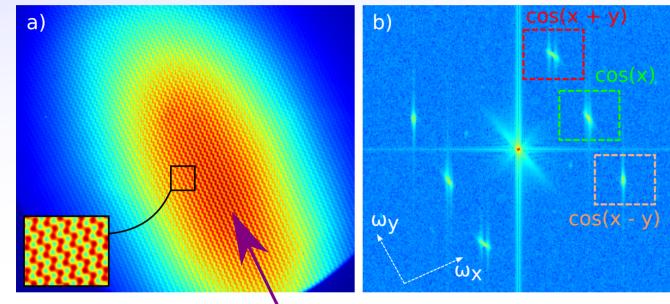


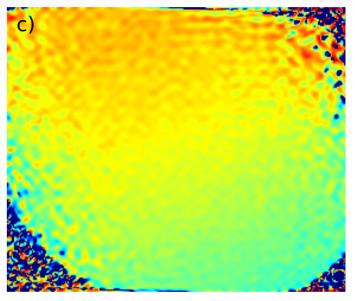
30 10 20 40 Polarisation θ / °

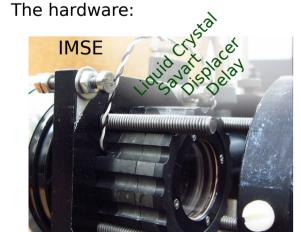


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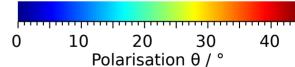
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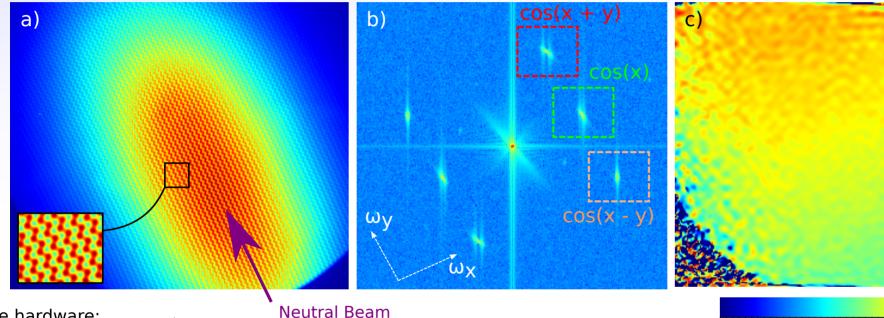
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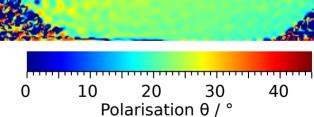
The hardware:

Neutral Beam

IMSE

Outstand And Anti-Set

Outstand Ant

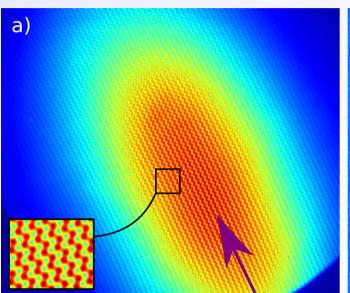


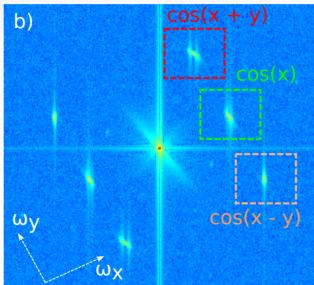


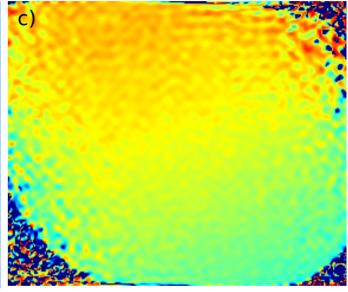


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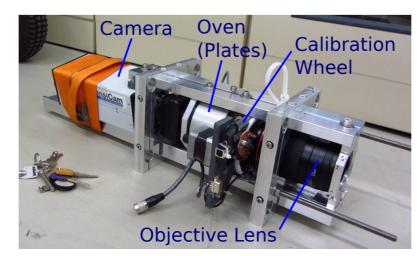


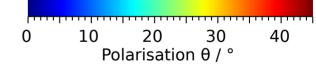


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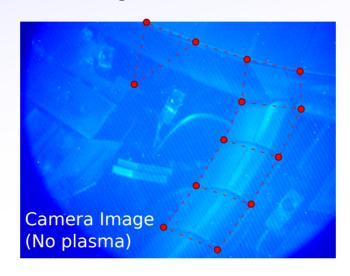


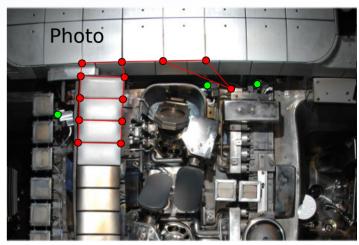


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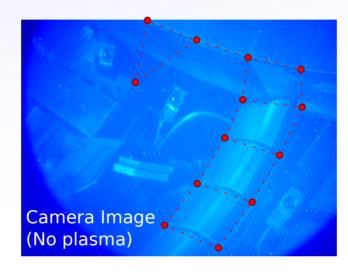


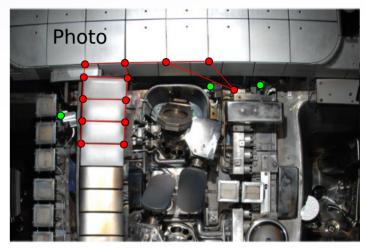


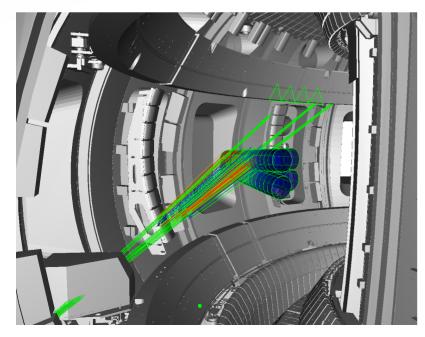


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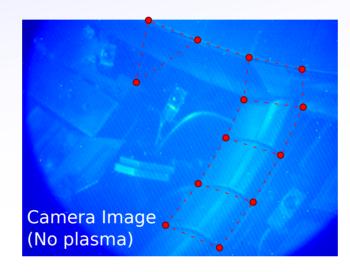


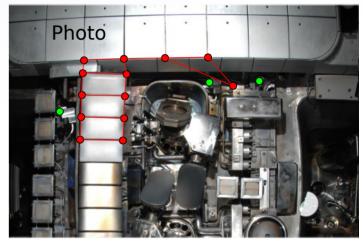


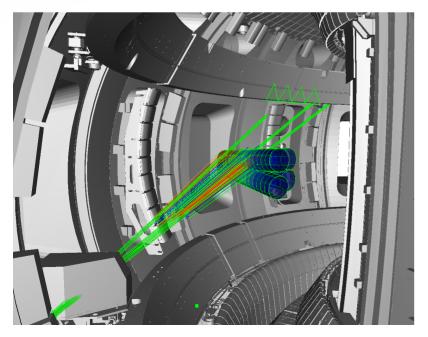


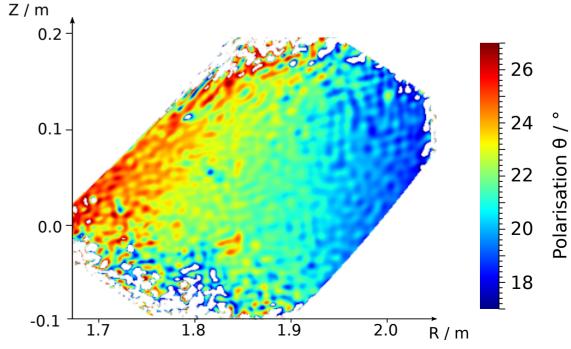
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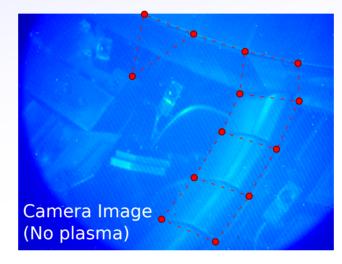


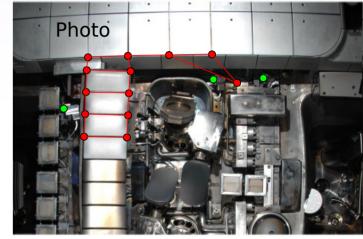


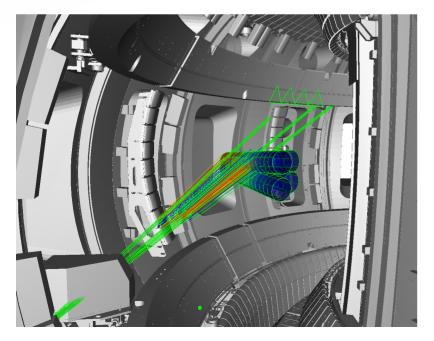
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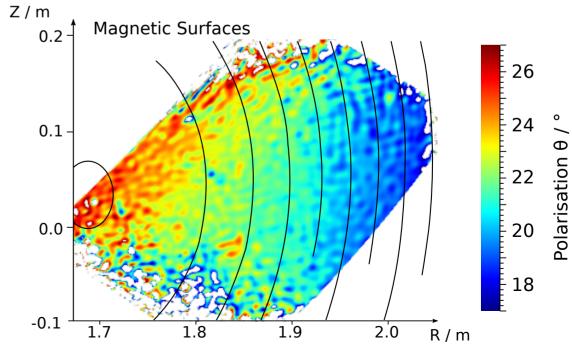
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The IMSE has required no torus access to calibrate (so far).











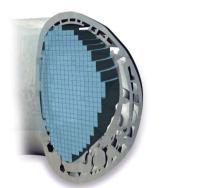


Analysis - Forward Modelling

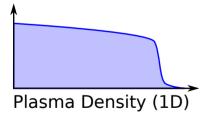
Videos are nice, but we also need to do some quantitive analysis of the data. The observed polarisation is a very complex function of the magnetic and electric fields, the neutral beam injection, the atomic/quantum physics of the Stark emission and the diagnostic optics.

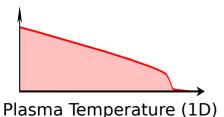


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Plasma Current (2D)

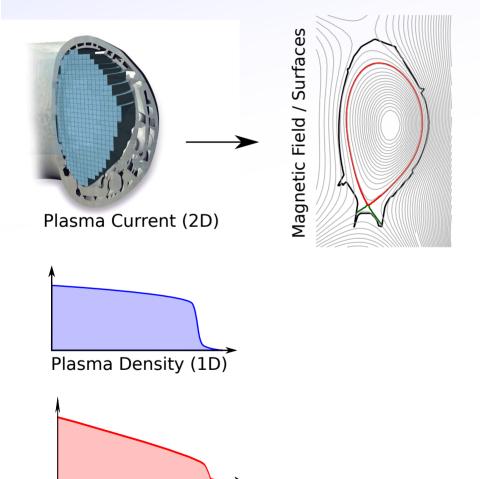






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We have a highly detailed modular forward modelling system for the entire plasma, to which we added components for the IMSE data:

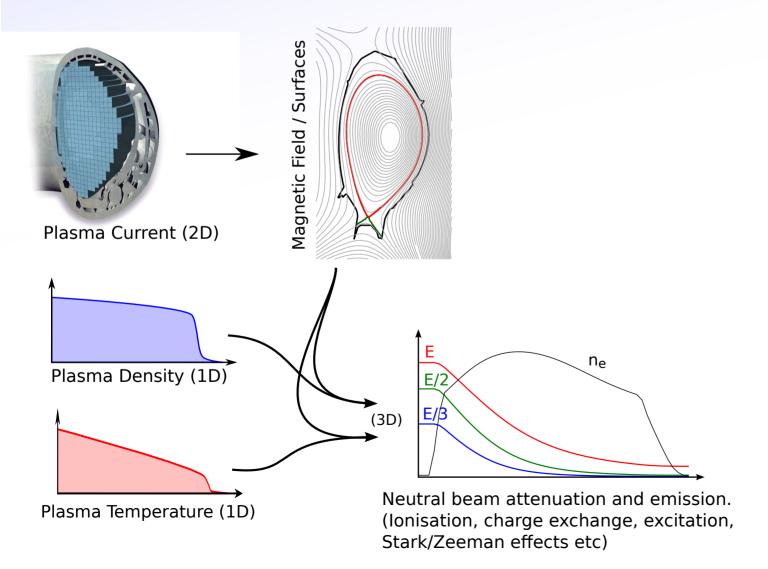


Plasma Temperature (1D)



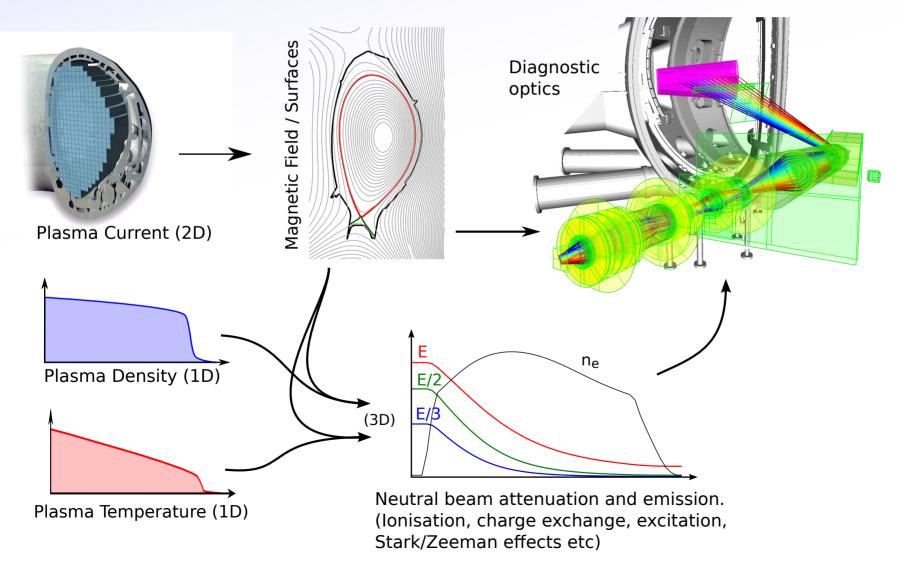


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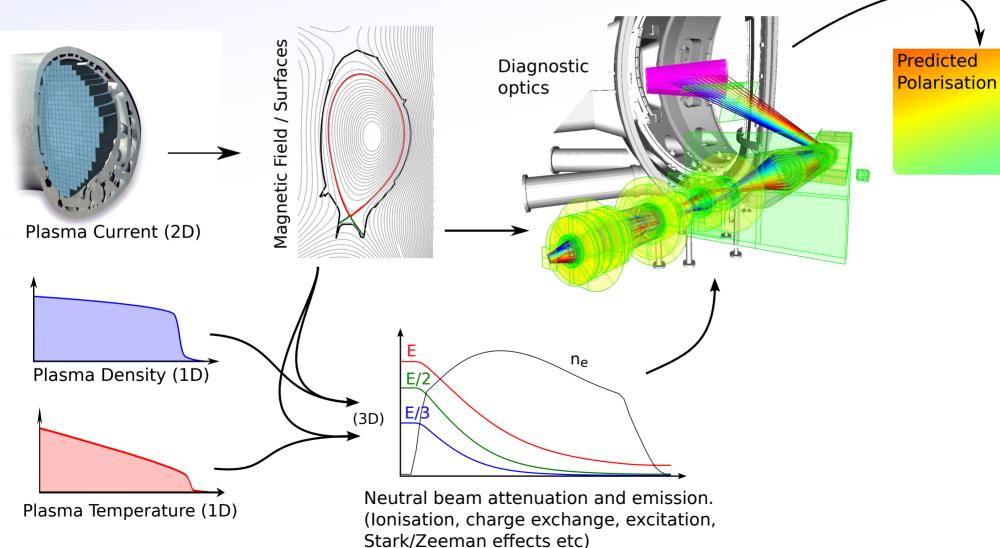
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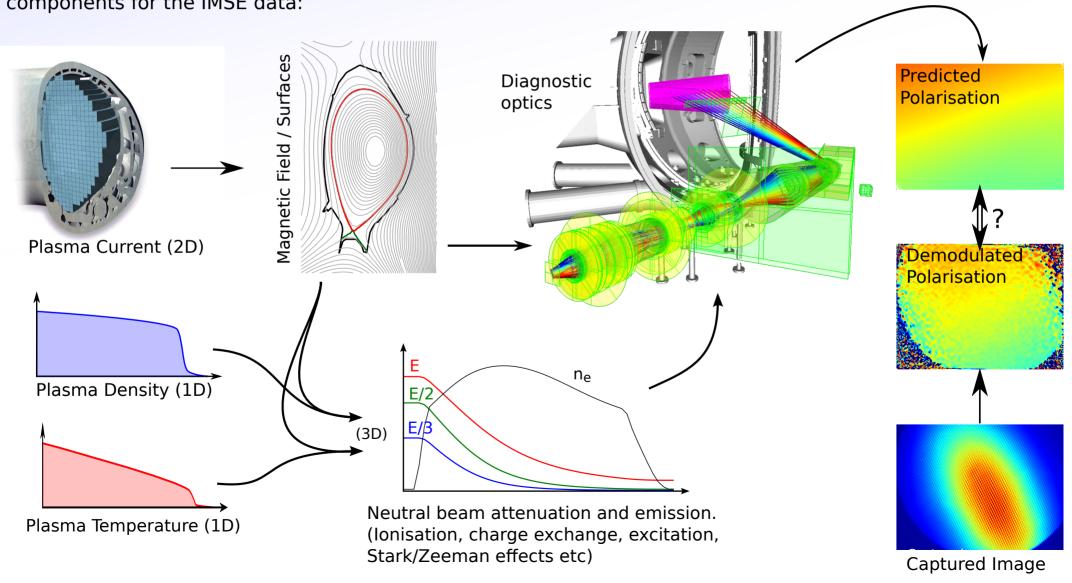
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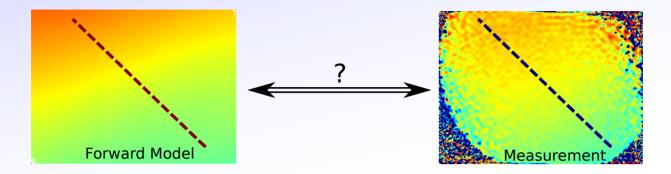


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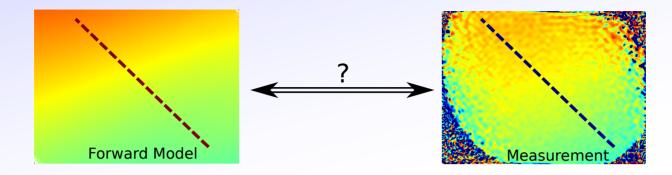




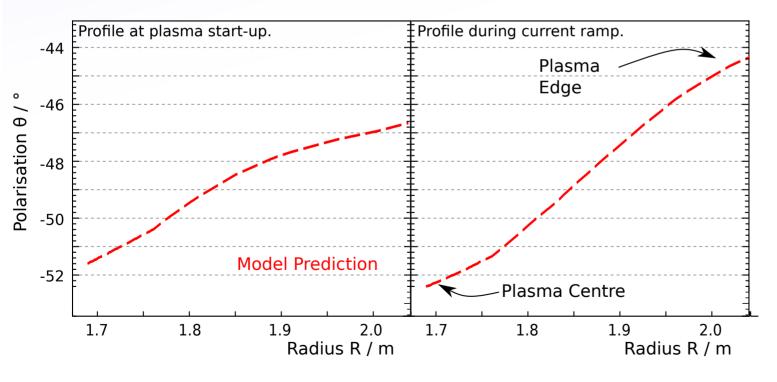






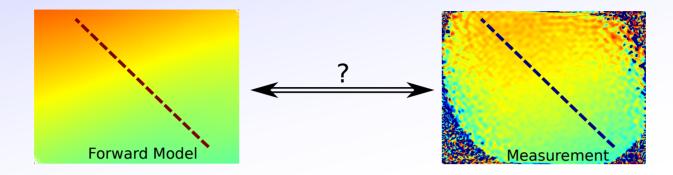


Except for a 0.7° offset, the results agree with the modelling where what we already know from other diagnostics is excepted to be correct. The difference is the new information that the IMSE provides:

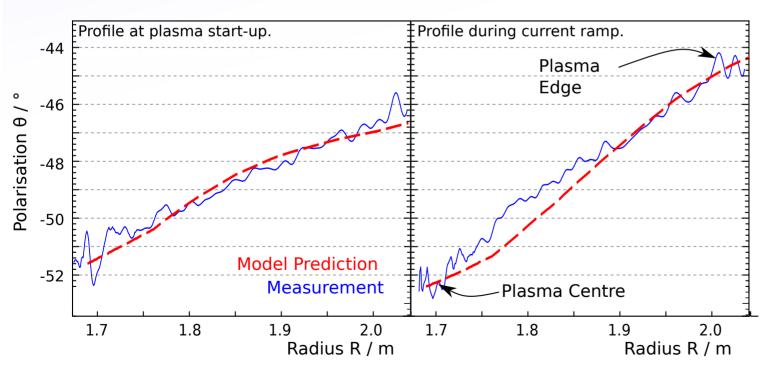






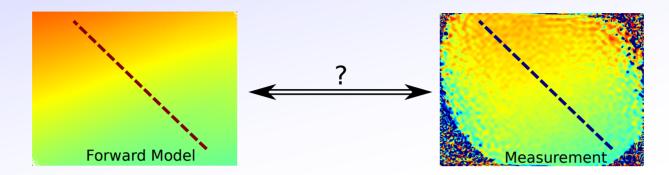


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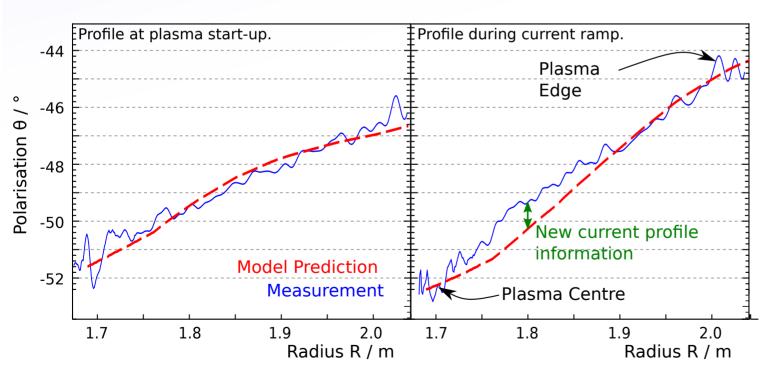






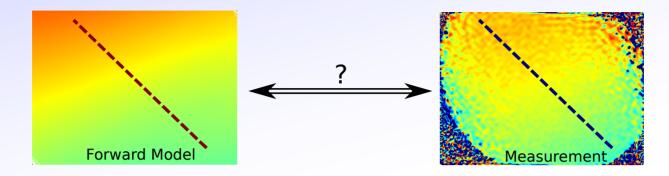


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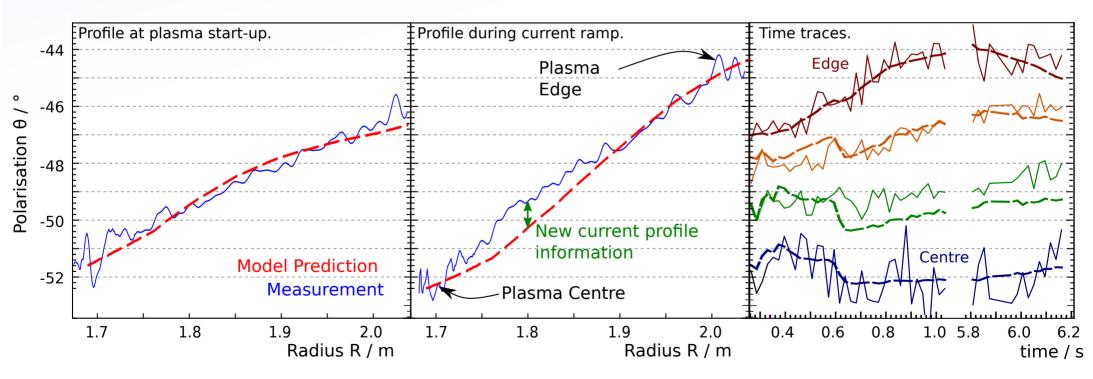








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Inferring the plasma current and q-profiles from the polarisation is far from trivial and the analysis work is still on-going.





Oliver Ford IPP Greifswald

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- An new Imaging Motional Stark Effect diagnostic has been designed, constructed and operated on ASDEX Upgrade.

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Over 100x more data.

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- Initial analysis shows agreement with modelled polarisation, within expected uncertainty.
- Next stage is to calculate safety factor profiles and plasma current image from the observed polarisation images.