

S42: NBI+ECRH in high-mirror

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Presented by Oliver Ford on behalf of the W7-X Team



TF-I Meeting. 3rd February 2023



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Proposals



Prio-I:		
oliford_002 stato_022	Threshold of P_ECRH into pure NBI for heat transport change ECRH into pure NBI with no X2 absoption	Scenario-development
dacar_006 thir_002	Turbulence in suppressed turb. scenarios (DR, PCI, CECE, SXR) Threshold P_ECRH for impurity transport change	Measurement specifics
alkn_004 cbra_011 kbr_10 mspolaor_002 tya_023	Beta effects on edge topology XMCTS Shafranov shift at different betas Alfvén Eigenmodes in high beta Edge EM turbulence in high beta (MPM probe) MHD stability in high performance	Measurements at high-beta

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F	Prio-2: astechow_011	Turbulence "matching" between pure ECRH and NBI	Modification request
F	Additional: cswee_001, twegner_007	Impurity transport in NBI+ECRH	Cover in passing



Anomalous heat flux oliford_002 - Threshold of P_{ECRH} into pure NBI for heat transport change 0.14 Hypothesis: Some threshold ECRH power exists where density pump-out High ECRH (±NBI) 0.12 is strong enough to switch to reduced turbulence. ÓO 0.10 \bigcirc 0 $[MWm^{-2}]$ 0.08 0.06 O anom 0.04 0.02 NBI and low ECRH 0.00 Pure NBI 10 12 $2n_e \nabla T_{i_{[10^{20}m^{-3}keVm^{-1}]}}$



oliford_002 - Threshold of P_{ECRH} into pure NBI for heat transport change0.14Hypothesis: Some threshold ECRH power exists where density pump-out
is strong enough to switch to reduced turbulence.0.12stato_022 - High ECRH into pure NBI with no X2 absoption
Hypothesis: Above $n_e > 1.5 \times 10^{20}$, only broad O2 deposition and no X2 absoption
so no large increase in turbulence, even with high O2 power.0.14









Wendelste



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- requires one pure-NBI shot with 3/4 sources for density profile check for O2. (also, energy extension?)





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For thir_002, additional specific request to test if peaking onset is supressed/delayed with low-level ECRH at start of pure-NBI. (Lower priority)





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3	NBI+O2, as #20221201.058 but lower ne	3MW	Q7, Q8	Trace He only, maybe GPI	stato_022, oliford_002, thir_002, +beta*		a *		
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5	NBI+O2, higher power (skip if Ti(#3) already drops to 1.5keV)	4MW	Q7, Q8	Trace He only, maybe GPI	stato_022, oliford_002	2, thir_002, +beta	a*		
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9	NBI+O2, identity #181009.034 above O2 cut-off	1MW	Q7, Q8	Trace He only, maybe GPI	stato_022, oliford_00	2, thir_002, +beta	1 *		
10	- ECRH only -		07.00	Trees Us only maybe CDI	-lifered 000 state 00	0 45 000 shate	*		
11	NBI+O2, threshold scan, between boundary in 3,5,7,9	1.5MW (tbd)	Q7, Q8	Trace He only, maybe GPI	oliford_002, stato_02	2, thir_002, +beta	1"		
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High-beta



Most prio-I proposals ask simply for 'high beta'. What we have achieved so far:

Shot	#181016.037	#181009.034	#221201.053	#221207.054	#221207.05	8
Short desc	Pellets HP	NBI+ECRH	NBI/2+ECRH	SS NBI+ECRH	Torsten's ham	imer
NBI		0	3.5	1.7	3.5	3.5
ECRH		5	1	1	2	3.5
ne dl [10^19]		8	11	9	5	16
Ne(0) [10^19]		8.5	12 7.5 (boz prelim)	5.5 (boz prelim)	25 (boz prelim)	
Wdia		1.1	0.7	0.5	0.5	1
Ti(0)		3	1.8	1.8	2	1.5
Te(0)		3.5	2	1.8	2	1.5
Axis beta	1.8	30%	1.50%	0.90%	0.70%	2.40%

With more sources / more gyrotrons, could be more...

Power supplies and Wdia



In #20221201.058 we achieved: 3.5MW NBI, 3.5MW ECRH, $n_e(0) \sim 2.2 \times 10^{20}$, $T_{i/e}(0) \sim 1.5$ keV, Wdia = 1.0MJ, beta(0) ~ 2.0%

We need to reduce the density to $\sim 1.8 \times 10^{20}$. Linear extrapolation --> Wdia = 900kJ 2 NBI sources + all existing gyrotrons available now (4.3MW) = 1.1MJ

If B1, F5 gyrotrons become available: Wdia = 1.4MJ 3 NBI sources and all gyrotrons: 6MW NBI + 3.9MW ECRH --> Wdia > 2MJ

Current ECRH status [component logbook page]:

- A1 820 250 kW -- shared power supply with NBI Q4 (1750 kW)
- A5 770 300 kW -- shared power supply with NBI Q8 (1750 kW)
- B1 550 kW (not yet fully available)
- B5 810 300 kW
- C1 600 250 kW -- shared power supply with NBI Q3 (1750 kW)
- C5 850 250 kW -- shared power supply with NBI Q7 (1750 kW)
- D1 710 250 kW
- D5 not available
- E1 790 400 kW
- E5 650 200 kW
- F5 550 kW (not yet fully available)

port

Plan:

I: Explore boundary between low and high heat diffusivity, optimise NBI/ECRH balance.

- Approach from higher and lower ECRH plasmas.
- Compare with similar 'optimum confinement' at W7-AS, drawing on experience. Is this similar physics?
- Compare global confinement of NBI/ECRH balance with similar series already performed at LHD.





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Analysis/results:

- Precise dependence of heat diffusivity on density gradient.
- Assesment of the generality and scaling of the scenario.
- An optimised NBI/ECRH scenario for use in later studies, e.g. impurity transport, turbulence, detachment compatibility etc...

With 4 NBI sources, *might* give access to up to $T_i = 2.5 keV$.





NBI + ECRH: Heat 1

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- Turbulent fluctuation measurements for in all scenarios for turbulence studies --> see talk by A. von Stechow





Wendelstei

Turbulent fluctuations from Doppler reflectometer reduced in NBI+low ECRH compared to equivalent ECRH



NBI + ECRH: Heat transport



Background:

- Steep density gradients lead to ~4x lower heat diffusivity than in high ECRH.
- Low ECRH power (< 1MW) does not significantly increase anomalous diffusivity, so allows increase in T_i .



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- Leads to very strong impurity accumulation.

Wendelstein

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